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Astronomy 104, Spring 2023

Test 2

CORRECT SOLUTIONS

Make sure your scantron has your name and code on it.

Show a picture ID, and turn in the test paper with the scantron.

Please circle the correct answers on the test paper.

There were many scrambled versions. Here is a solved copy of one of the versions.

- **1** Depends on the version.
- **2** Depends on the version.
- **3** Depends on the version.

B Atoms consist of the following constituents:

- A: Electrons, protons and neutrons.
- B: Electrons, and a nucleus.
- C: Quarks, electrons, and gluons.
- D: Electrons and nucleons.
- E: Molecules and electrons.

C How was the speed of light measured for the first time?

- A: By measuring the slowing down of the time in a spaceship.
- B: By measuring the delay in conversations of astronauts.
- C: Measuring the delay in the eclipses of Jupiter's moons.
- D: Getting a laser reflection from the Moon.
- E: By direct measurement in a laboratory experiment.

C How do we know the chemical composition of stars?

- A: From samples returned by spacecraft.
- B: It is calculated based on the amount of energy the star radiates.
- C: From the presence of each atom's spectral lines.
- D: From a chemical analysis of interstellar gas blown in into the solar system.
- E: From a chemical analysis of cosmic rays.

D How do we measure the mass of stars?

- A: We measure the strength of their magnetic field.
- B: We measure how strong a gravitational effect they have on the motion of Earth.
- C: Using Wien's law, applied to the color of the star.
- D: Using Kepler's III law for binary stars.
- E: Using Kepler's II law applied on their planets.

A How long does light take to arrive from the Sun to Earth?

- A: 9 minutes.
- B: 1 light year.
- C: an hour.
- D: 4 years.
- E: 4.5 billion years.

C The energy of each photon is determined by ...

- A: the speed of the light.
- B: nothing at all: it is a universal quantum constant of nature.
- C: its wavelength only.
- D: the strength of the light only.
- E: both the wavelength and the strength of the light.

D The equivalent of light at much shorter wavelength is called ...

- A: Neutrinos.
- B: Radio waves.
- C: Charged particle radiation.
- D: X-rays.
- E: Ultraviolet.

E The spectrum of a hot glowing body is ...

- A: emission spectrum.
- B: band spectrum.
- C: a mixture of an emission and an absorption spectrum.
- D: absorption spectrum.
- E: continuous spectrum.

D What can you tell from the wavelength of a spectral line?

- A: The strength of gravity of the source object.
- B: The temperature of the gas that emits the light.
- C: The atmospheric pressure in the source object.
- D: Which atom/molecule produced the line.
- E: How far is the source from the observer.

B What does ionization mean?

- A: The breakup of molecules into atoms.
- B: Atoms losing (some of) their electrons.
- C: A chemical reaction with an ionizing salt.
- D: Mixing interstellar gas with Fe-containing dust.
- E: Electrons moving to a higher orbit in an atom.

E What happens to matter at 10,000 degrees?

- A: All substances become solid.
- B: Atomic nuclei change into each other.
- C: All atoms become ionized.
- D: Molecules break up.
- E: (At least some) atoms break up.

B What is plasma?

- A: Gas that contains no electrons.
- B: (At least partially) ionized gas.
- C: Gas that contains no free electrons.
- D: A viscous liquid.
- E: Hot gas with most molecules broken up into atoms.

C Where do 1. low energy, 2. high energy cosmic rays come from?

- A: All: beyond the Solar System
- B: All: the Solar Wind
- C: Low: the Solar Wind, High: deep space.
- D: All: from Earth, mainly nuclear reactors
- E: Low: deep space, High: the Solar Wind.

E A hydrogen nucleus ...

- A: is a neutron.
- B: is a proton and an electron.
- C: consists of two protons.
- D: consists of one proton and one neutron.
- E: is a proton.

A At what temperature will all molecules disintegrate?

- A: Above ~ 2000 K
- B: Above ~ 10,000 K.
- C: Above ~ 273 K
- D: Above ~ 1 million K
- E: Above ~ 100 K

E How does the Doppler-effect affect the spectrum of a star?

A: The color of a moving star looks redder/bluer than normal.

B: The spectral lines of an approaching star are shifted from the red end of the spectrum to the blue end.

C: The star's light is stronger when the star is approaching us, and weaker when it is receding.

D: An approaching star's light arrive to Earth sooner, which causes its spectral lines broaden.

E: Spectral lines shift (usually a tiny bit) when the star moves towards or away from us.

E How does the speed of light relate to the speed of radio waves and the speed of sound?

A: Sound and radio waves have the same speed, the speed of light is infinitely fast.

B: Radio waves and sound are slow, light is fast.

C: They all have the same speed.

D: Radio waves are fastest, light is middle, sound is slowest.

E: Radio waves and light have the same speed, sound is much slower.

A What can excite a hydrogen atom?

- A: UV radiation.
- B: Heating to 1200 K.

C: Red light.

D: Infrared radiation.

E: A magnetic field.

E What type of a spectrum does the Sun have?

- A: emission spectrum.
- B: continuous spectrum.
- C: band spectrum.
- D: a mixture of an emission and an absorption spectrum.
- E: absorption spectrum.

B What type of spectrum does fluorescence produce?

- A: Band spectrum.
- B: Emission spectrum.
- C: Distorted spectrum.
- D: Absorption spectrum.
- E: Continous spectrum.

C Which one is an example of the Doppler-effect?

- A: Magnetic fields make spectral lines break up into several lines.
- B: Gas illuminated by UV radiation glows in an emission spectrum.
- C: The spectral lines of a moving star are shifted.
- D: The overall color of a moving star changes.
- E: The motion of a star makes spectral lines broaden.

C How large is the Doppler effect in astronomy in practice?

A: It is impossible to detect the Doppler effect in the spectra of individual stars.

B: Both the position of the spectral lines and the overall color of the star changes noticeably.

- C: A tiny (say, 0.01%) shift in the wavelength of spectral lines.
- D: A large shift (say, 10%) in the wavelength of spectral lines.
- E: A large shift in the overall color of a star, say, from blue to red appearence.

C Relate the energy of a blue photon to the energy of a red photon.

- A: Blue photons have half the energy of red photons.
- B: Blue photons have a hundred times less energy than red photons.
- C: Blue photons have twice as much energy as red photons.
- D: Cannot tell: the energy of photons depends on the strength of the light.
- E: Blue photons have a hundred times as much energy as red photons.

D What happens to the spectrum of a star that rotates?

- A: It is shifted towards red or blue by an even amount.
- B: Spectral lines become more prominent.
- C: The spectral lines in the red part of the spectrum become brighter.
- D: All its spectral lines broaden the same way.
- E: They are shifted by an amount proportional to their wavelength.

A The wavelength of light determines its ...

- A: color.
- B: speed.
- C: ratio of electric / magnetic field strength.
- D: intensity.
- E: polarization.

B What change occurs to molecules in hot matter such as the matter of a star? At what temperature?

- A: Molecules are fused into very large atoms at ~ 1200 degrees.
- B: Molecules break up into atoms at ~ 2000 to 3000 degrees.
- C: Molecules break up into atoms at ~ 2 million degrees.
- D: Molecules turn into protons and electrons at ~ 2000 degrees.
- E: Molecules pick up extra electrons at ~ 10,000 degrees to become ions.

E When do you see an absorption spectrum?

- A: When charged particle hit magnetic fields.
- B: When planets reflect sunlight.
- C: When hot objects glow.
- D: When gas is illuminated from the side.
- E: When cold gas is in front of a light source.

A Which of the following is wrong?

- A: The stronger the light, the more energy each of its photons has.
- B: X-ray photons have more energy than photons of light.
- C: A blue photon always has more energy than a red photon.
- D: Microwave photons have less energy than photons of light.
- E: The energy of each photon does not depend on how strong the light is.

E Which one is moving fastest?

- A: Visible light
- B: Radio waves
- C: Laser
- D: X-rays
- E: The same

C Why do we use spacecraft to do X-ray astronomy?

- A: Because spacecraft is closer to the stars.
- B: Because X-ray telescopes radiate dangerous amounts of X-rays.
- C: Because the atmosphere is opaque in X-rays.
- D: Because the telescope needs to be cooled to liquid helium temperatures (3K).
- E: Because of interference from medical use of X-rays.

E Why does a planetary nebula glow?

- A: It is illumnated by the Sun.
- B: It reflects the light of the star in its center.
- C: It is illuminated by the planet whose atmosphere it is.
- D: It is illuminated by surrounding stars.
- E: It fluoresces in the UV of the central white dwarf.

D What chemical process changes carbon into oxygen?

- A: Hydrolysis.
- B: Burning.
- C: Oxydation.
- D: None.
- E: Electrolysis.

B A star's color indicates its ...

- A: chemical composition.
- B: temperature.
- C: mass.
- D: size.
- E: distance.

B What does Stefan-Boltzman's law say?

A: Each absorption line is brightest at a particular temperature.

B: The power radiated by a surface element of a glowing body is proportional to the 4th power of its temperature.

C: The temperature of a star is inversely proportional to the wavelength of its color.

D: The absolute magnitude of a star is independent of its distance.

E: The speed of rotation of a star is inversely proportial to its temperature.

B What does Wien's law say?

A: Hotter objects appear redder (I.e. radiate in longer wavelength).

B: Thermal glow gets of shorter wavelength when the temperature of the body is increased.

C: The wavelength of the light of an approaching body is shifted towards blue.

D: Hotter gas radiates stronger than cold, proportionally to the fourth power of temperture.

E: Warmer material absorbs red light stronger.

C What is a double-line binary star?

A: A star that is visibly broken up into a pair in the telescope and the orbital motion is noticeable.

B: A doule star system with strong stellar wind.

C: A short-period binary with both stars' spectral lines visible.

D: A double star with a strong magnetic field.

E: A star that is visibly broken up into a pair in the telescope but the orbital motion is not noticeable.

B Why are double stars important in astronomy?

- A: It is easier to detect a double star than a single one.
- B: They provide the only way to tell stellar masses.
- C: Double stars have many planets orbiting around them.
- D: They provide the only way to determine the chemical composition of stars.
- E: They are the only ones that blow up as supernovae.

C The spectral type of the Sun is ...

- A: M
- B: B
- C: G
- D: K
- E: A

A What percentage of all stars are doubles?

- A: Half.
- B: All.
- C: There are no true double stars. They may seem so only due to perspective.
- D: Almost none.
- E: Almost all.

E What spectral type is the red giant Betelgeuse?

- A: C.
- B: G.
- C: Hydrogen.
- D: He.
- E: M.

B The spectral type of a star is related to ...

- A: the temperature in its core.
- B: its surface temperture.
- C: its distance.
- D: its speed of motion.
- E: its chemical composition.

B Strong ultraviolet radiation comes only from stars of spectral type ..., and why?

- A: All spectral types, because they are all hot.
- B: O & B, because these are hot enough.
- C: K & M, because these are hot enough.
- D: All spectral types, because UV production does not depend on temperature.
- E: None, because stars do not radiate in UV.