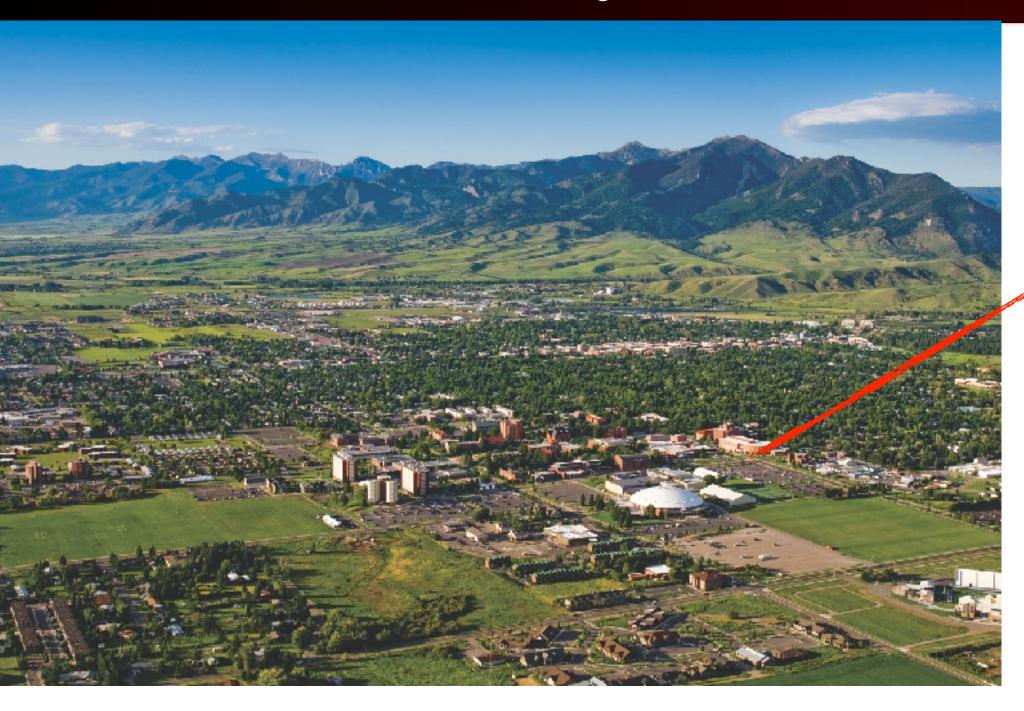
## Future Tests of General Relativity with Gravitational Waves

Barausse, Yunes, Chamberlain, arXiv 1603.04075, PRL (2016) Yunes, Yagi, Pretorius, arXiv 1608.06187, PRD (2016) Chamberlain and Yunes, to be submitted (very soon!)

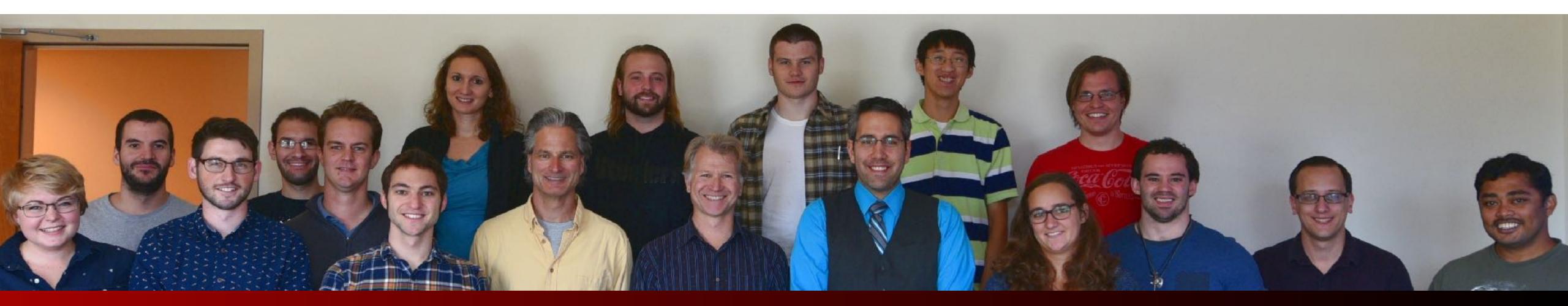
Nicolas Yunes eXtreme Gravity Institute Montana State University

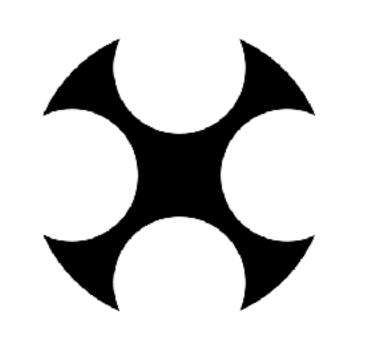
StronG BaD Workshop March 2nd, 2017

### eXtreme Gravity Institute









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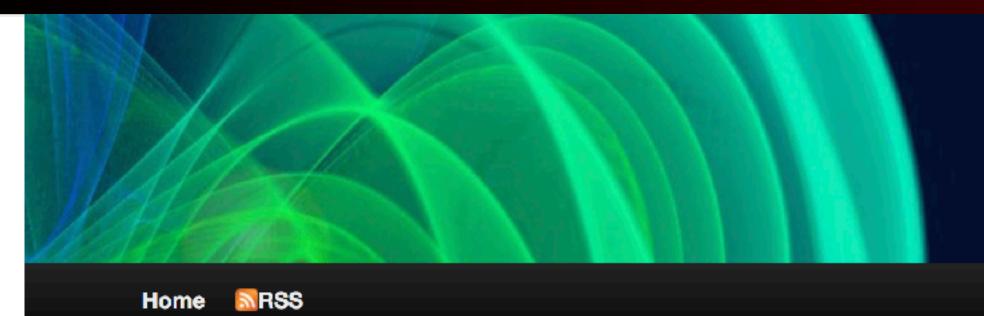








## eXtreme Matter meets eXtreme Gravity



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**GRG Scientists** 

More info: external link Contact: ncornish[AT]montana.edu Date: 2017-08-17 - 2017-08-19 Location: Bozeman, Montana, USA

#### eXtreme Matter meets eXtreme Gravity Workshop, Bozeman, Montana, USA

XGI Workshop First Announcement:

"eXtreme Matter meets eXtreme Gravity" August 17-19, Bozeman Montana

The eXtreme Gravity Institute at Montana State University will hold a workshop to discuss methods for constraining the properties of Neutron Stars and the dense-matter equation of state. Like previous XGI workshops, the format will emphasize discussion and exchange of ideas over formal presentations. Each session will be organized around a science question, with a moderator and two discussion leaders. Topics to be covered include gravitational-wave observations of Neutron Star – Neutron Star and Neutron Star – Black Hole binaries, X-ray observations by the NICER mission (set to launch very soon), theoretical calculations of the dense-matter equation of state, and numerical simulations of NS-NS and NS-BH mergers.

The meeting is being held immediately prior to the HEAD meeting in Sun Valley, and participants may choose to drive between the meetings, or simply head a little south of Bozeman to view the total eclipse on the 21st of August. Bozeman is a beautiful mountain town a one-hour drive from the North entrance of Yellowstone National Park. The surrounding area offers great opportunities for hiking, fishing, white water rafting, and mountain biking.

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Editor: Luciano Rezzolla

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#### CONFERENCES

- ICRANet-Minsk workshop on high energy astrophysics, Minsk, Belarus
- Fifth Galileo-Xu Guangqi Meeting, Chengdu, China
- 15th Italian-Korean Symposium on Relativistic Astrophysics, Seoul, Korea
- Geometric Foundations of Gravity in Tartu, Estonia
- 3rd Karl Schwarzschild Meeting Gravity and the Gauge/Gravity Correspondence, Frankfurt, Germany

#### JOBS

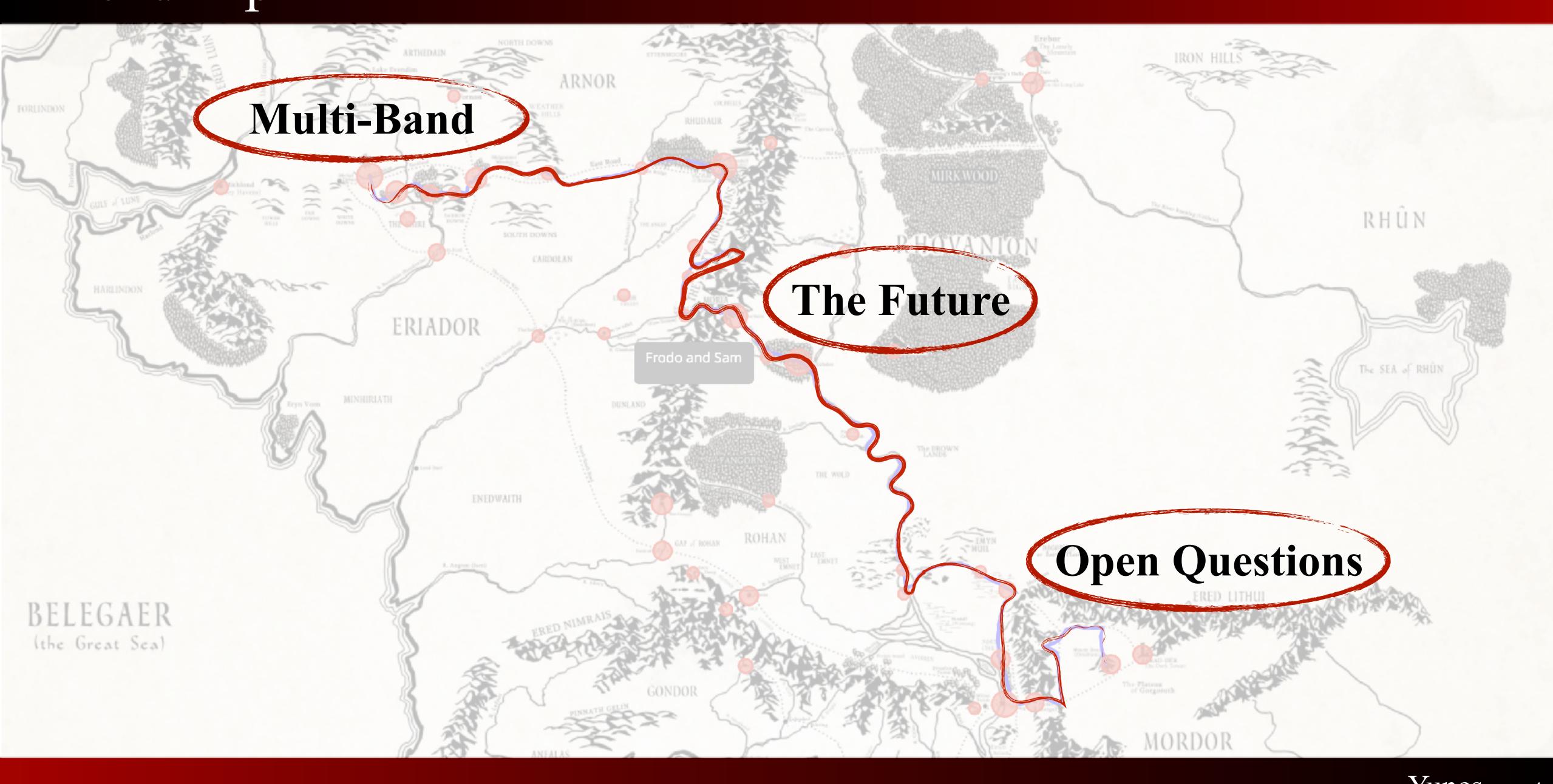
- Assistant Lecturer in Gravitational Wave
   Astrophysics at Monash University, Australia
- Professor/Reader in Gravitational Wave Science at Portsmouth LIK





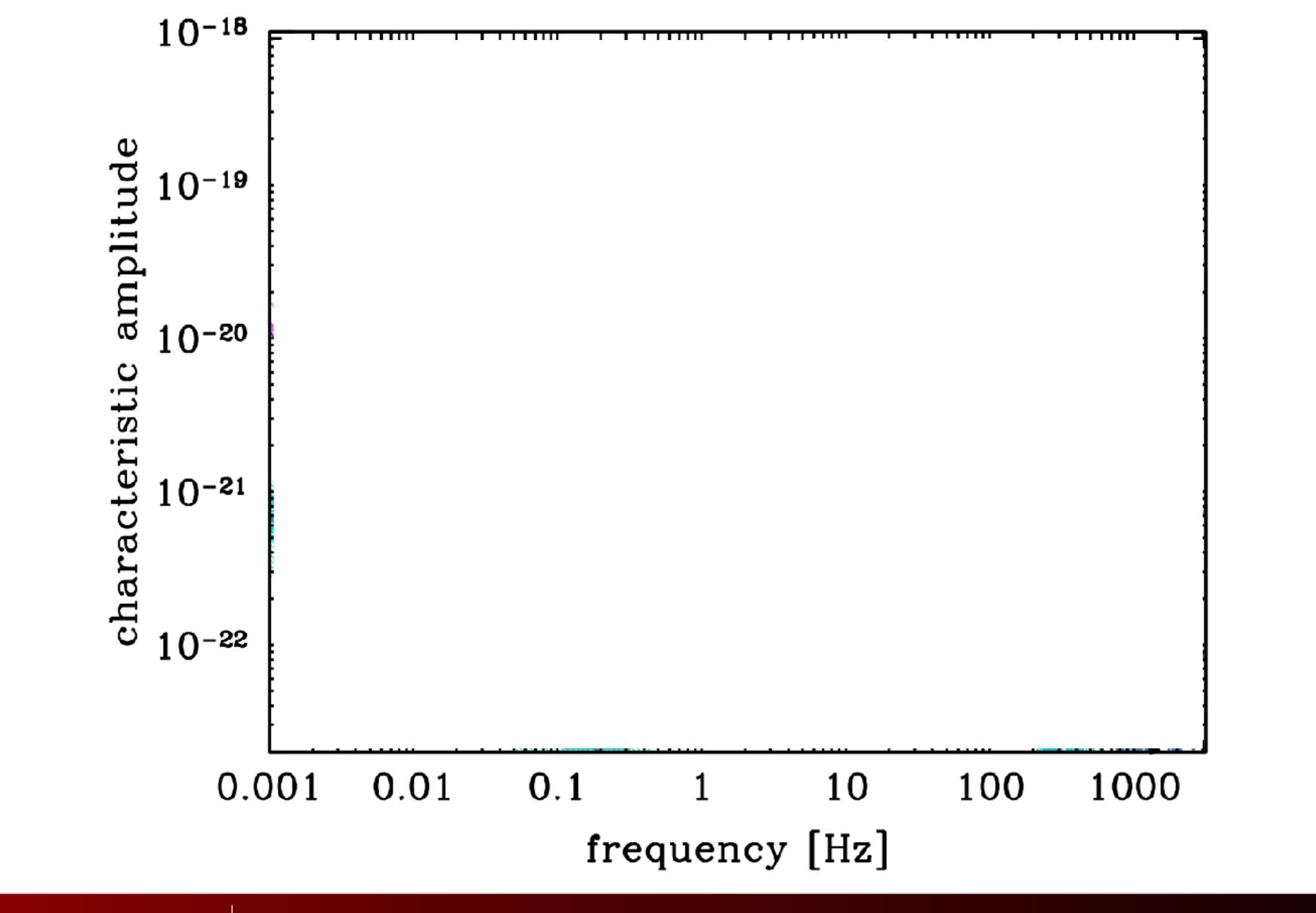


### Roadmap





### Multi-Band Observations



Multi-Band

The Future

**Open** Questions

[Amaro-Seoane & Santamaria Ap.J. '10, Sesana, PRL '16]



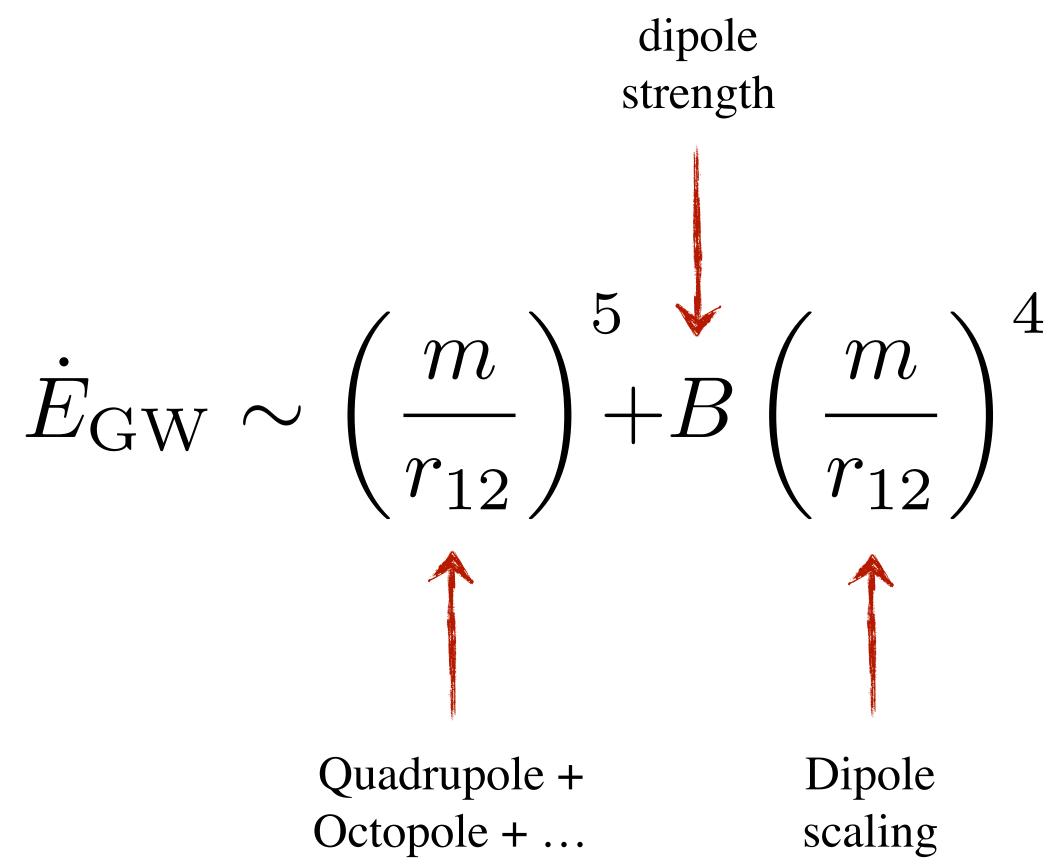




## Dipole Energy Flux

Quadrupole +

Octopole + ...









### Effect of Dipole Flux

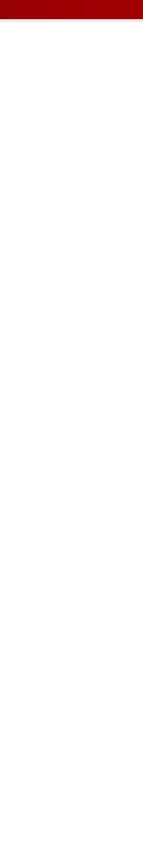
### Faster Inspiral, Faster Merger

$$\frac{df_{\rm GW}}{dt} = \left(\frac{df_{\rm GW}}{dE_{\rm b}}\right) \left(\frac{dE_{\rm b}}{dt}\right) = -\left(\frac{df_{\rm GW}}{dE_{\rm b}}\right) \dot{E}_{\rm GW} = -\left(\frac{df_{\rm GW}}{dE_{\rm b}}\right) \left[\left(\frac{m}{r_{12}}\right)^5 + B\left(\frac{m}{r_{12}}\right)^4\right]$$

### Faster Inspiral, Faster Gravitational Wave Chirp

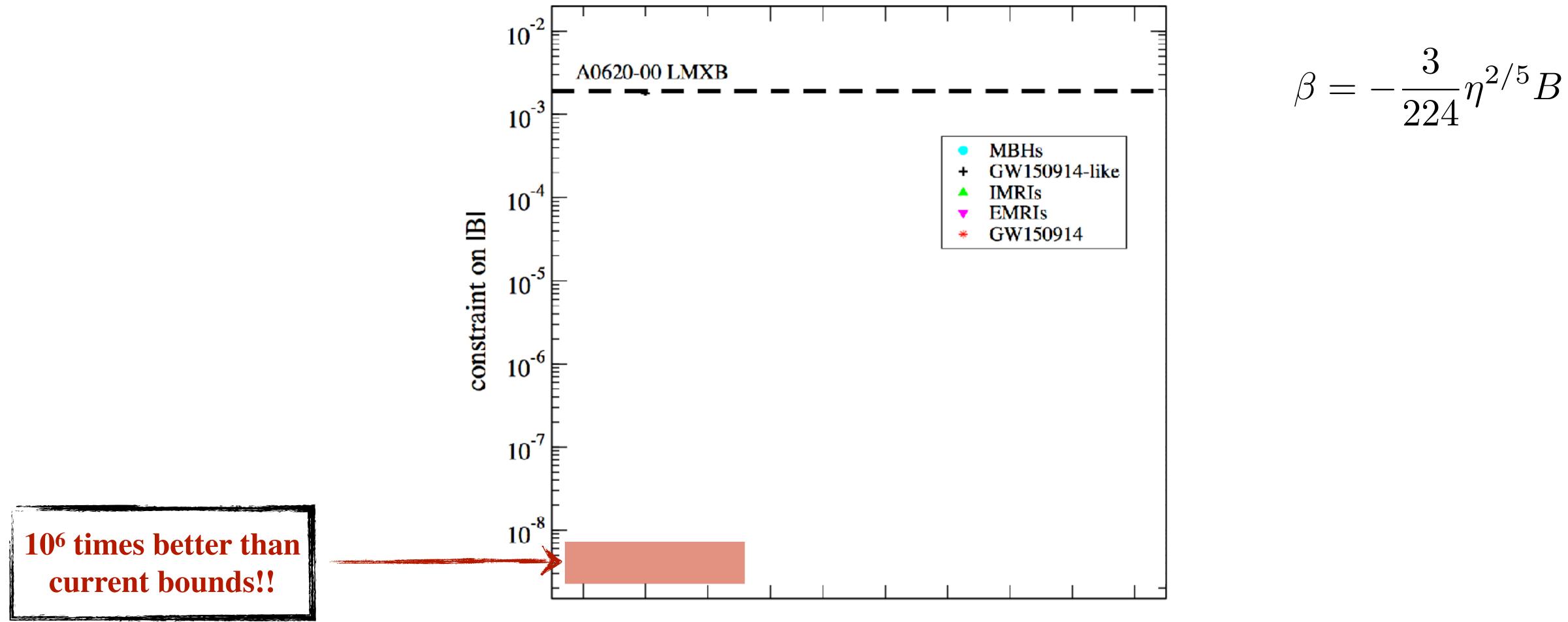
$$\Psi_{\rm GW} \propto \int \int \int t' \left( \frac{df_{\rm GW}}{dt''} \right) dt'' dt' = \Psi_{\rm GW,GR} + \beta(B) \left( \pi \mathcal{M} f_{\rm GW} \right)^{-7/3}$$
Quadrupole + Dipole
Octopole + ... GW Term







## Projected Constraints on Dipole Radiation



.

[Barausse, Yunes, Chamberlain, PRL '16]

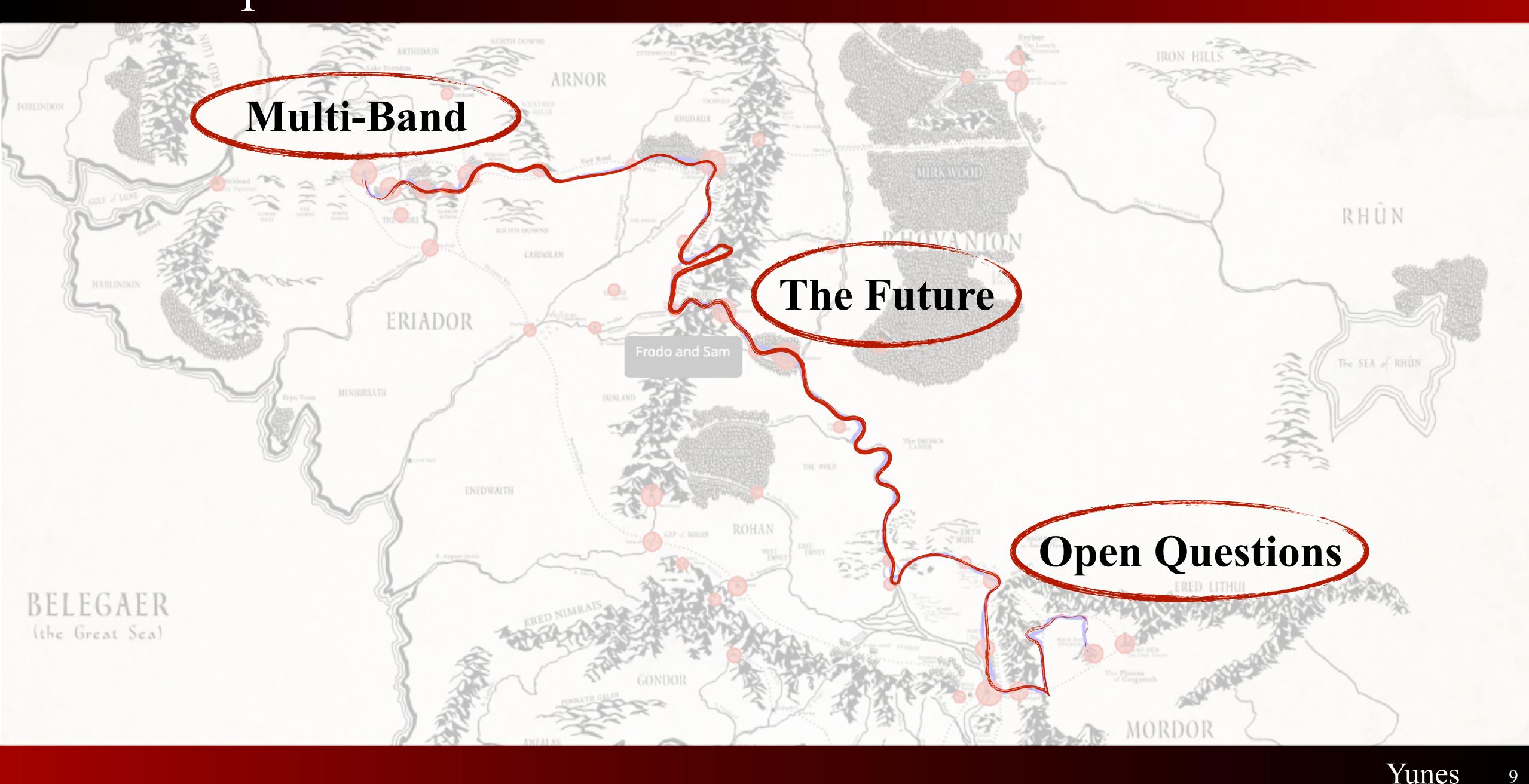




Yunes



### Roadmap





# Estúpidamente Preliminary Greek for "somewhat"

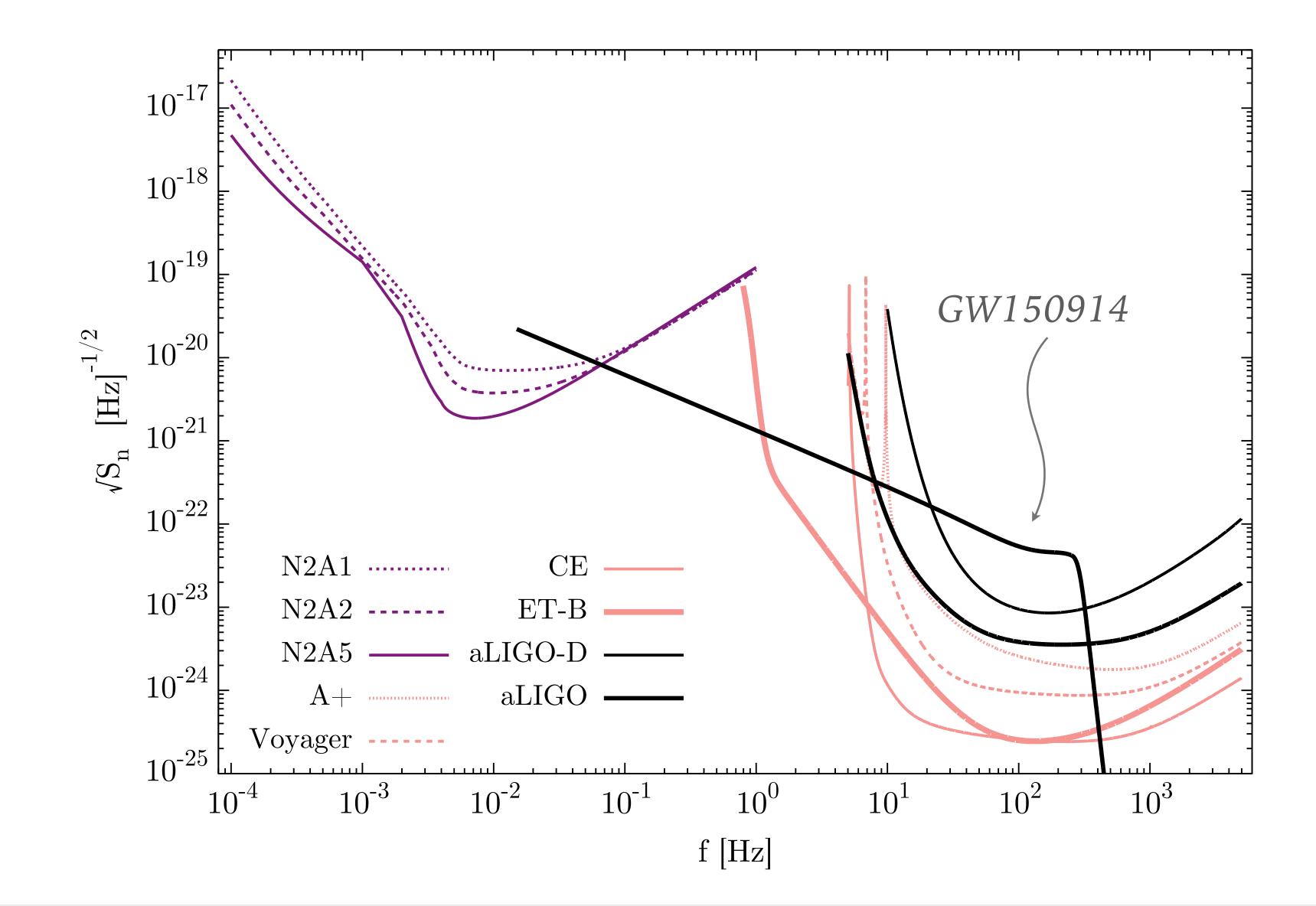
The Future **Open** Questions Multi-Band







### Spectral Noises of Future Instruments



Multi-Band

#### The Future

#### **Open** Questions





## In life you have to make choices, so we made some...

_							
	Name	$m_1[M_\odot]$	$m_2[M_\odot]$	$(\chi_1,\chi_2)$	$D_L$	$\mathbf{Z}$	ho -rang
-	GW150914	35.1	<b>29.5</b>	(0.31, 0.39)	400 Mpc	$\sim 0.09$	6 - 12
	EMRI	$10^5$	10	(0.8, 0.4)	1 Gpc	$\sim 0.2$	60 - 23
	IMRI	$10^{5}$	10 <sup>3</sup>	(0.7, 0.9)	$5~{ m Gpc}$	$\sim 0.8$	166 - 65
	IMBH	$5 \times 10^3$	$4 \times 10^3$	(0.7, 0.9)	$16~{ m Gpc}$	$\sim 2$	58-22
_	$\mathbf{SMBH}$	$5 \times 10^{6}$	$4 \times 10^{6}$	(0.7, 0.9)	$48  \mathrm{Gpc}$	$\sim 5$	372 - 11

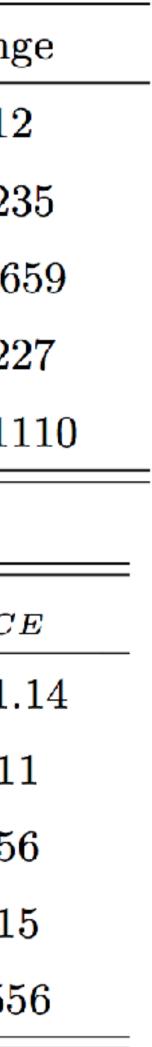
Name	$m_1[M_\odot]$	$m_2[M_{\odot}]$	$(\chi_1,\chi_2)$	$D_L$	$\rho_{aLIGO}$	$ ho_{A+}$	$ ho_{Voyager}$	$ ho_{ET-B}$	$\rho_{CH}$
NSNS	2	1.4	(0.01, 0.02)	$100 { m Mpc}$	23.23	33.65	109.55	332.98	541.
$\ell \mathrm{BHNS}$	5	1.4	(0.2, 0.02)	$150 { m ~Mpc}$	22	31	103	312	511
ℓBHBH	8	5	(0.2, 0.3)	$250 { m ~Mpc}$	28	40	132	398	656
BHBH	25	20	(0.3, 0.4)	$800 { m Mpc}$	26	38	123	372	615
GW150914	35.1	29.5	(0.31, 0.39)	$400 { m Mpc}$	66	95	310	951	155

Multi-Band

#### The Future

#### **Open** Questions

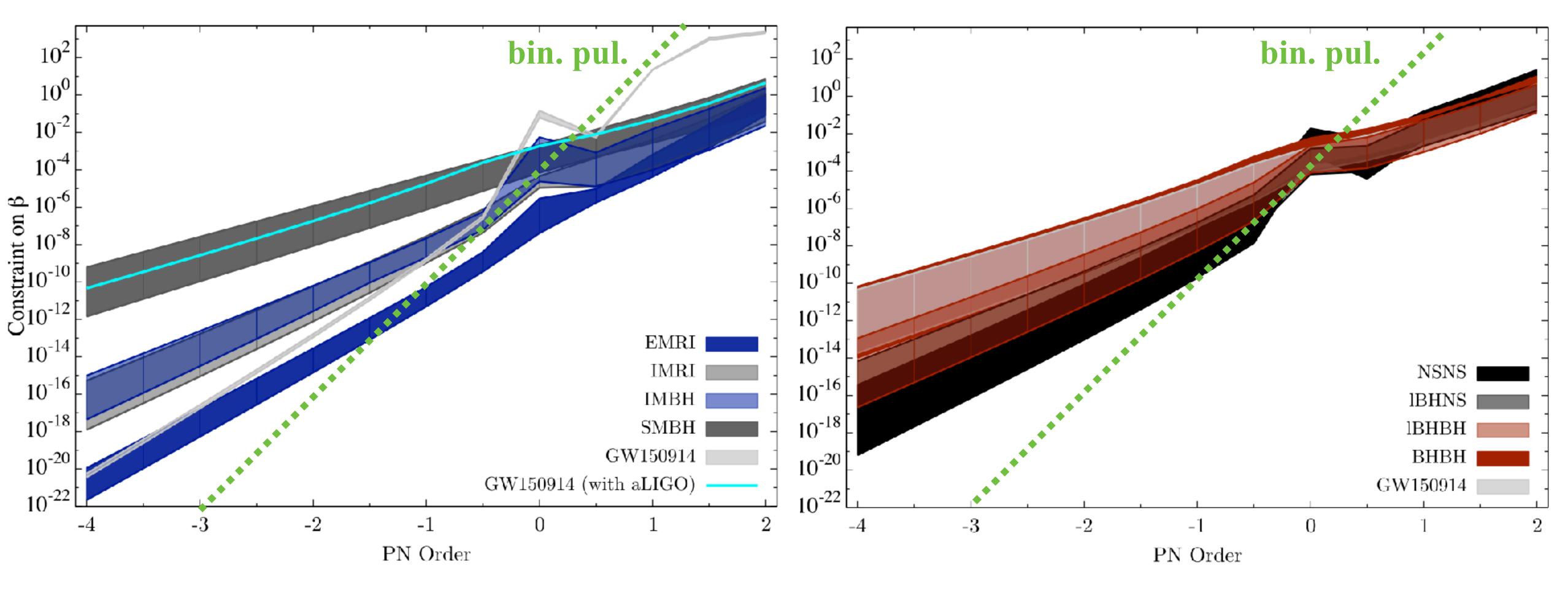






### Future ppE Constraints on GR

space-based



Multi-Band The Future **Open** Questions

#### ground-based

[Chamberlain & Yunes, to appear soon]

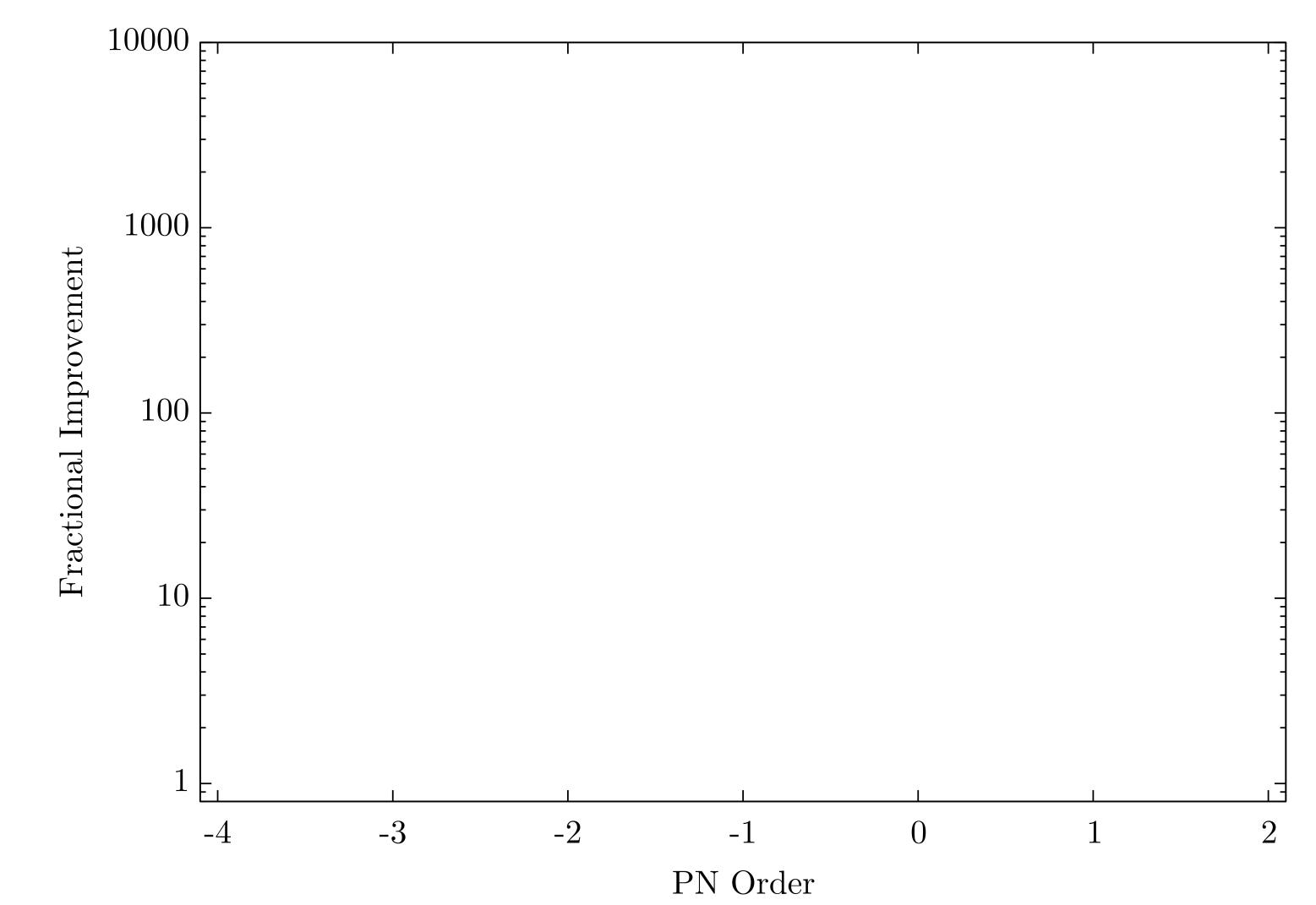








## Fractional Improvement of ppE Constraints



#### Multi-Band

#### The Future

#### **Open** Questions

[Chamberlain & Yunes, to appear soon]









### What are we really learning with GWs?

**Open** Questions The Future

Multi-Band

- **Violations of the Strong Equivalence Principle** 
  - **Lorentz Violations in Gravity**
  - **Gravitational Parity Violation**
- What matters the most is the *mapping* between **ppE constraints and theoretical physics inferences** 
  - **Graviton Mass and Propagation Effects** 
    - •
- (leaving out a lot of stuff here, e.g. no-hair tests with ringdown)







## Future Constraints on the Graviton Mass



Binary systems that are as far away as possible (SMBHs)

Binary with largest chirp mass

### **Open Questions:**

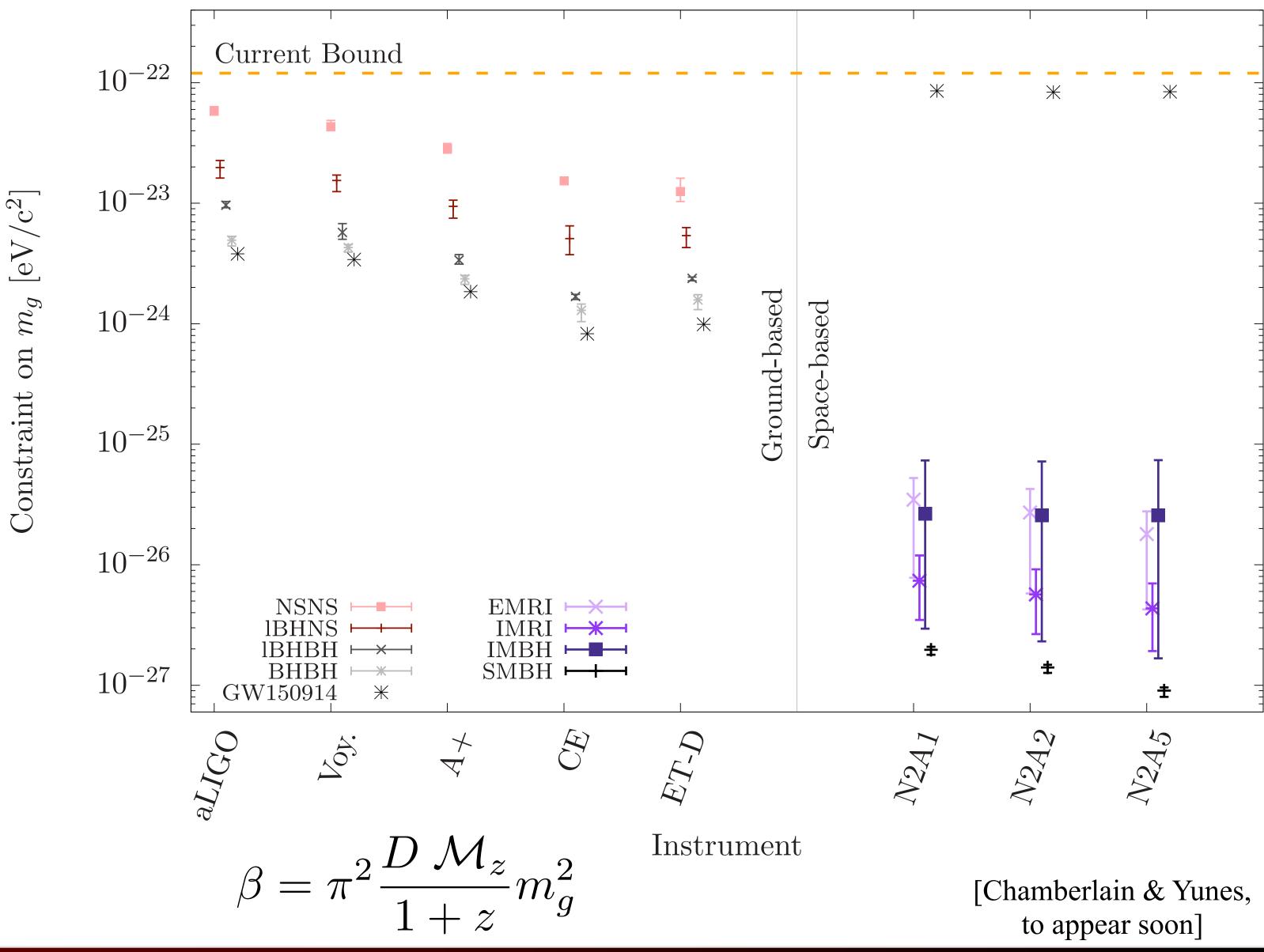
Multi-Band

Generation of GWs?

Merger? Hybrid IMR waveforms?

The Future





**Open** Questions





## Future Constraints on Violations of SEP

### **Extractable Physics:**

Non-Schw BHs (yes-hair theorem in EdGB)

NSs have scalar charge (scalar-tensor)

Compact Object binaries inspiral faster due to dipole radiation

### **Maximize Extraction:**

Low-mass BH or NS (long-inspiral) GWs

Binary with tiny mass ratio

### **Open Questions:**

Merger? Hybrid IMR waveforms?

 $10^{-3}$  $10^{-4}$  $10^{-5}$  $10^{-6}$  $10^{-7}$ 

 $\delta \dot{E}$ 

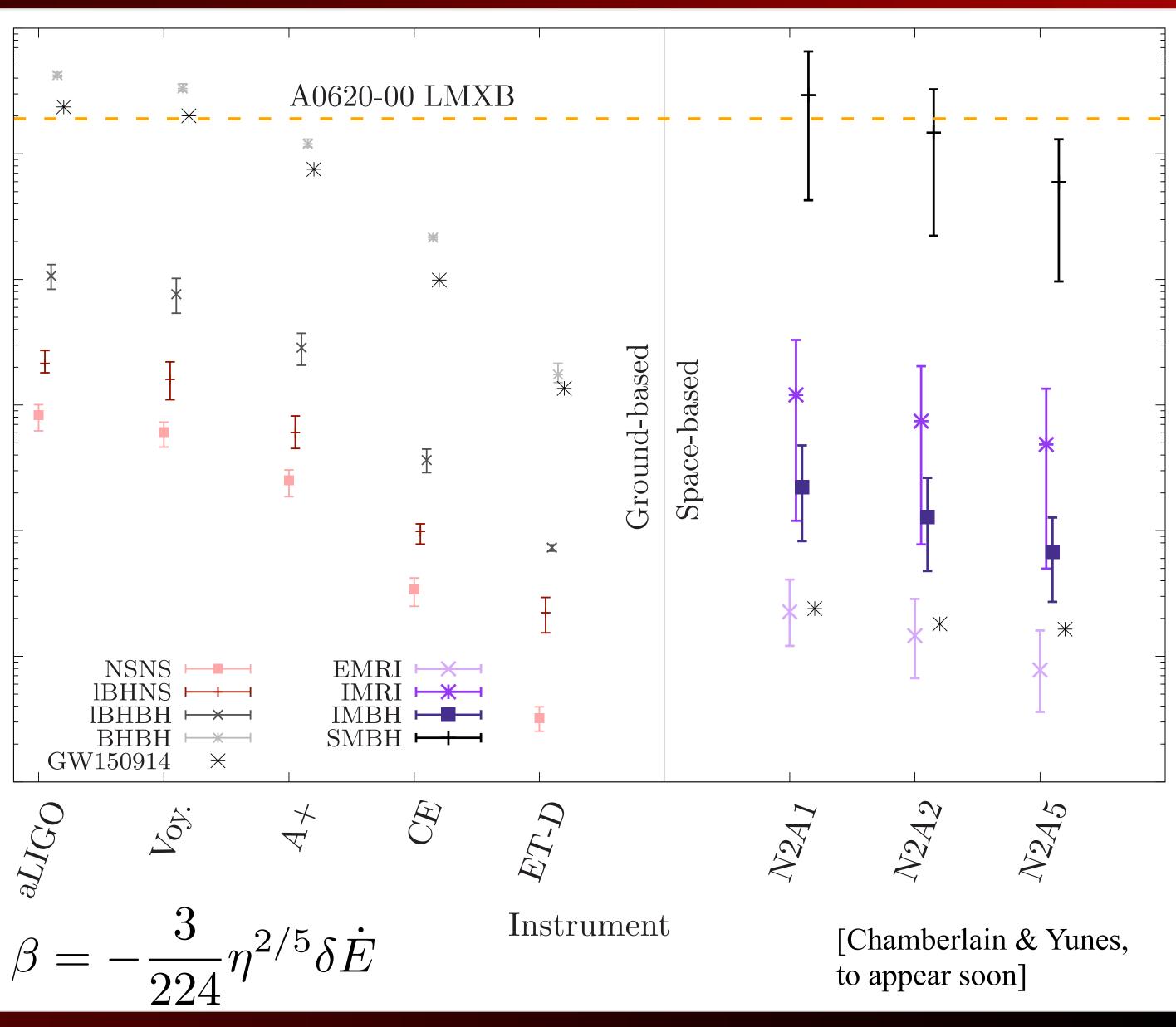
Constraint on

 $10^{-2}$ 

 $10^{-8}$ 

#### Multi-Band

#### **Open** Questions The Future







## Future Constraints on Gravitational Lorentz Violation

C

**Open** Questions

0.1

### **Extractable Physics:**

Non-Spinning BH is not Schwarzschild NSs have sensitivity-dependent GR deviations 0.01 Compact Object binaries inspiral faster (dipole

### **Maximize Extraction:**

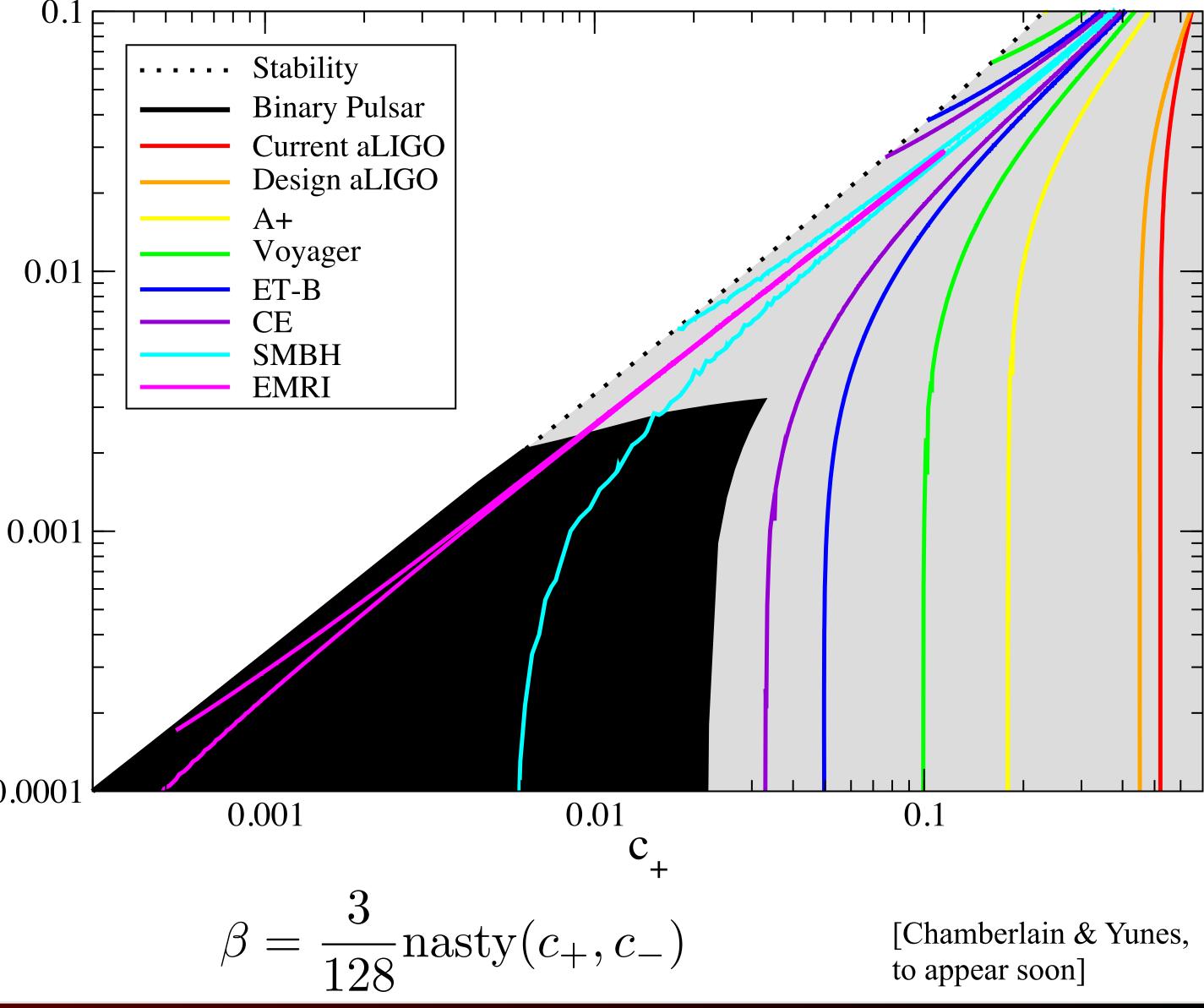
SMBHs or EMRIs do best

### **Open Questions:**

Multi-Band

BH sensitivities and Inspiral BH waveforms? 0.000Merger? Hybrid IMR waveforms?

The Future









## Future Constraints on the Variation of Newton's G



Binary system at widest separation possible (lowest frequency)

Binary with largest chirp mass

### **Open Questions:**

Generation of GWs?

Merger? Hybrid IMR waveforms?

$$\beta = -\frac{25}{65526} \frac{\dot{G}}{G} \mathcal{M}_z$$

[Chamberlain & Yunes, to appear soon]

$$10^{-4}$$
 -  $10^{-6}$  -  $10^{-8}$  -  $10^{-10}$  -  $10^{-12}$  -  $10^{-14}$  -  $10^{-16}$  -  $10^{-1$ 

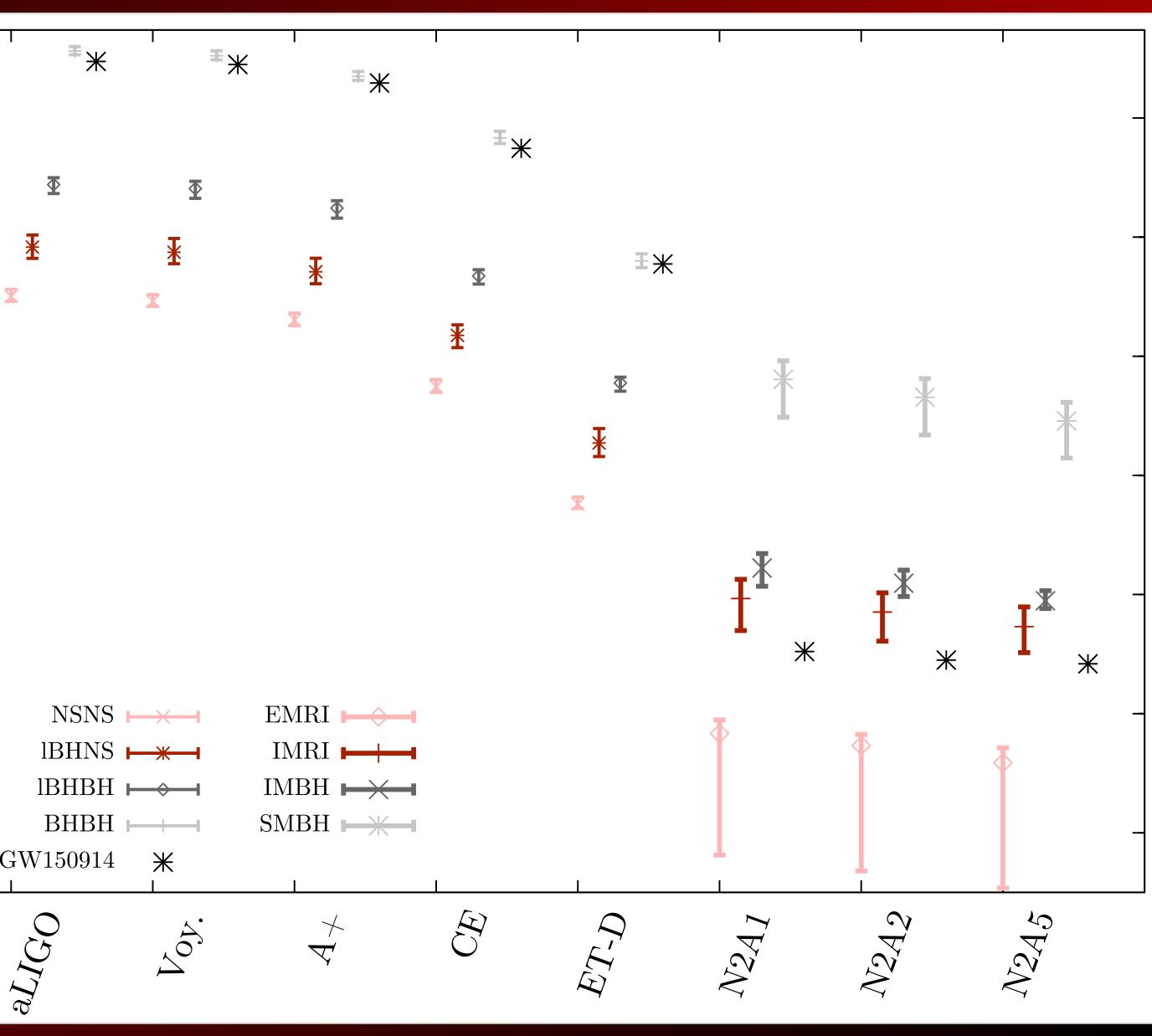
Gdot/G

Constraint on

Multi-Band

The Future

**Open** Questions

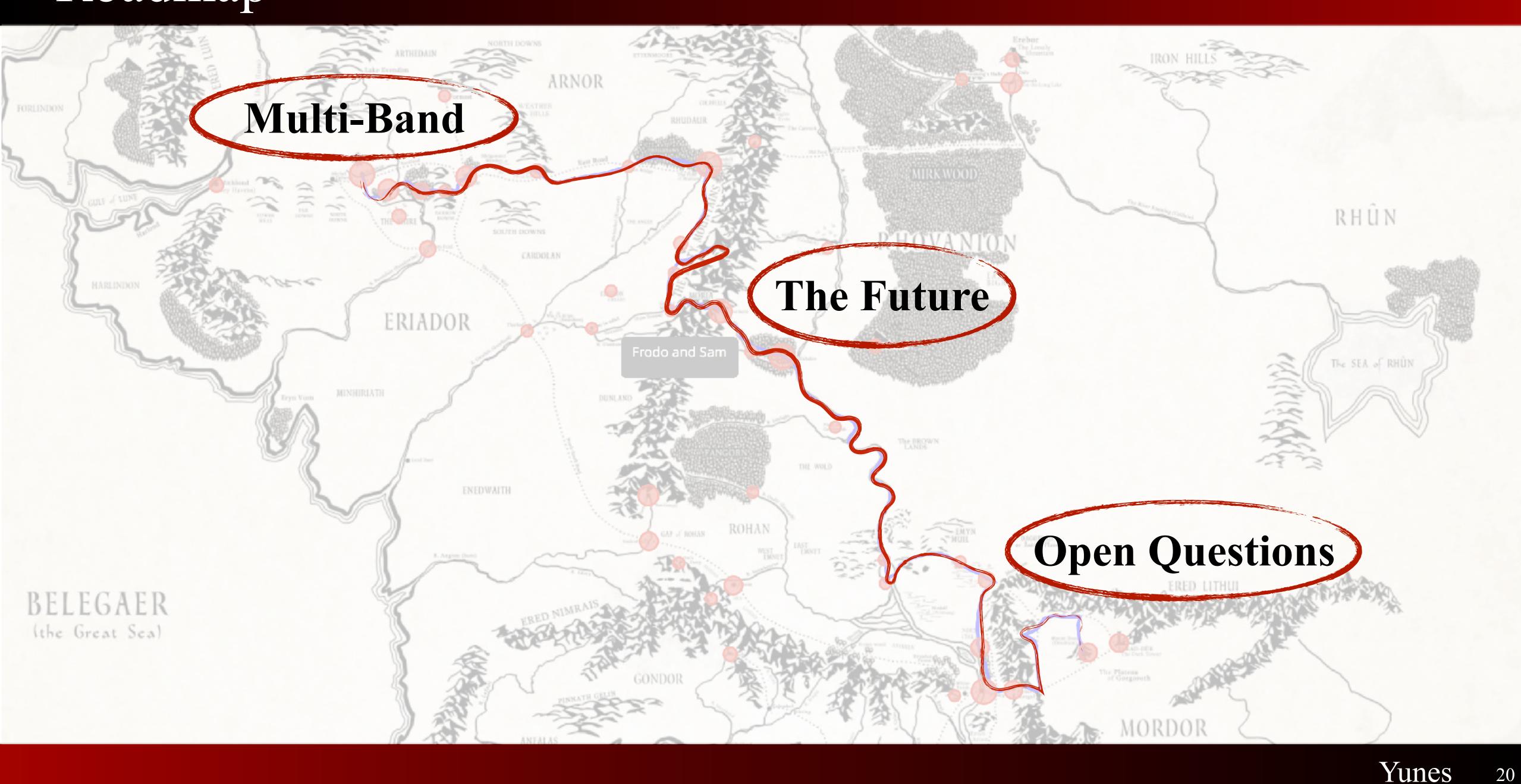








### Roadmap





### And now what?



### Duty

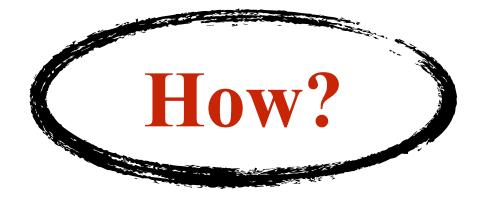
### Curiosity

### Religion

Multi-Band

The Future

**Open Questions** 



### **More Consistency ?**

### **Better ppE ?**

### **Compelling Alternative?**







### Nico's Crystal Ball

#### Observations: 25/year **5** years

<u>Detectors</u>: aLIGO —> A+, AVirgo, Kagra

<u>*Theory*</u>: in GR: IMR of precessing inspirals and (maybe) eccentric in not GR: mergers and 1 PN in a few theories (e.g. EdGB)

#### **10 years** Observations: 100/year

<u>Detectors</u>: A+ —> Voyager, AVirgo, Kagra, Indian LIGO <u>Theory</u>: in GR: single model for spin-precessing, eccentric, with NR calibration in not GR: mergers in more theories with spin precession

15 years **Observations:** rate limited

Multi-Band

<u>Detectors</u>: Voyager —> CE, aVirgo, aKagra, Indian LIGO, LISA <u>*Theory*</u>: in GR: EMRIs (with second-order self-force) in not GR: EMRIs in modified gravity?

**Open** Questions The Future









### Some Open Problems

### Theory

New & Interesting Physical Mechanisms? Cosmological Modified Theories?

Spin Precession in Modified Gravity?Mergers in Modified Gravity?EMRIs and resonances in Modified Gravity?

"New" sources of GWs? (eg. eccentric)

Multi-Band

The Future

Open Questions

### Experiment

Efficient data analysis w/spin precession? Reduced Order Methods for Mod. Grav. ?

Improved instruments vs new instruments? Combining EM information ?

Ringdown tests? Stacking?





## Thank You









## Future Constraints on Gravitational Parity Violation

**Extractable Physics:** 

Left/Right polarized waves propagate differently Spinning BH is not Kerr (yes-hair theorem) Spinning BH binaries inspiral faster (magnetic dipole)

## **Maximize Extraction:**

Multi-Band

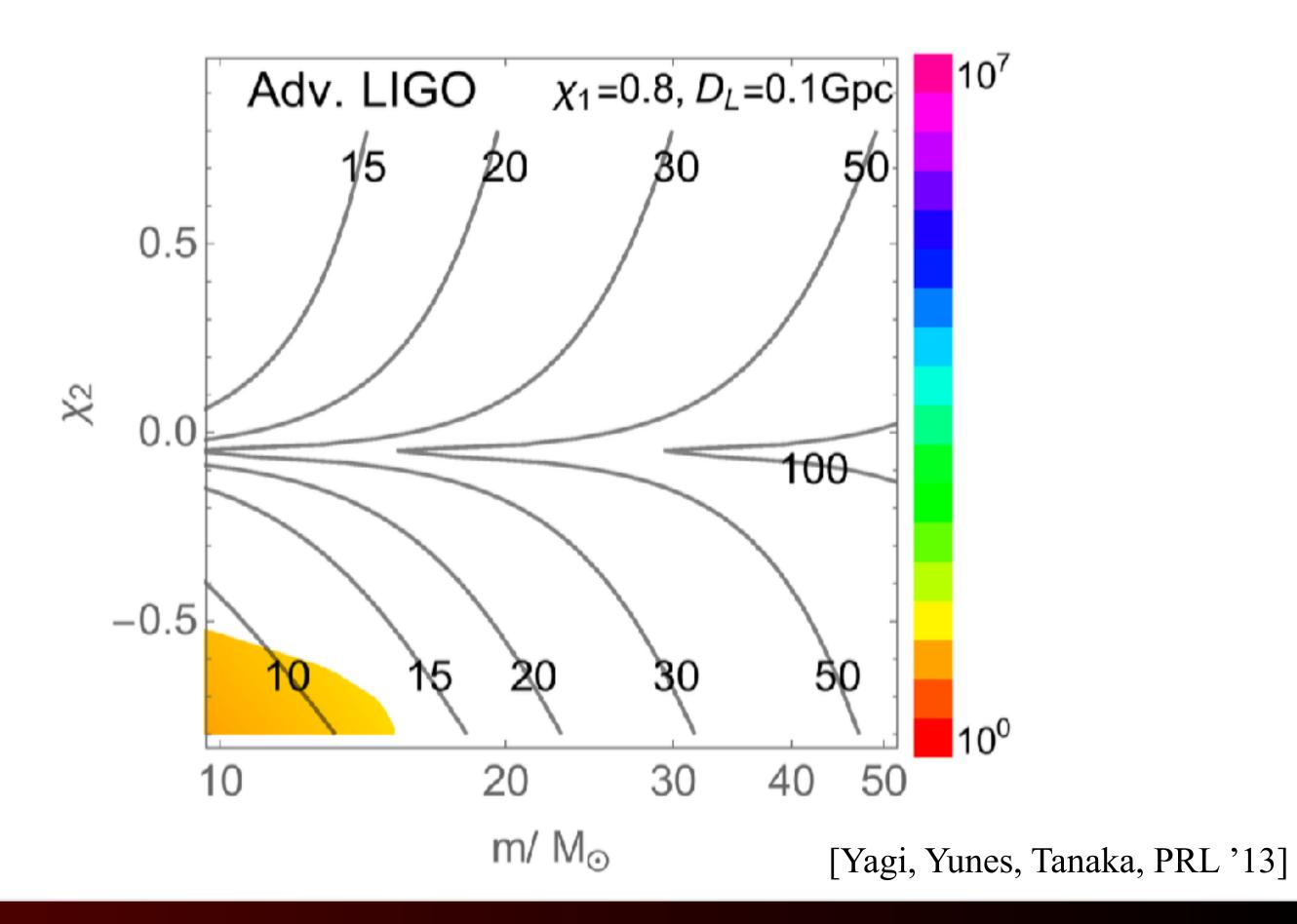
Low-mass (long-inspiral) high spin (precession) BH binaries

**Open** Questions

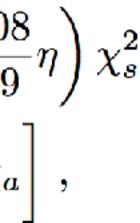
**Open Questions:** Merger? Hybrid IMR waveforms?

The Future

 $\beta_{\rm dCS} = \frac{1549225}{11812864} \frac{\zeta_{\rm dCS}}{\eta^{14/5}} \left[ \left( 1 - \frac{231808}{61969} \eta \right) \chi_s^2 \right]$  $+\left(1-rac{16068}{61969}\eta
ight)\chi_{a}^{2}-2\delta_{m}\chi_{s}\chi_{a}
ight|\,,$ 











### Parametrized post-Einsteinian Framework

### **The parameterized post-Einsteinian Framework**

$$\tilde{h}(f) = \tilde{h}_{GR}(f) \left(1 + \alpha f^a\right) e^{i\beta f^b}$$

Theoretical Effect	Theoretical Mechanism	Theories	ppE b	Order	Mapping
Scalar Dipolar Radiation	Scalar Monopole Field Activation	EdGB [140, 142, 149, 150]		-1PN	$\beta_{ m EdGB}$ [140]
Scalar Dipolar Radiation	BH Hair Growth	Scalar-Tensor Theories [59, 151]	-7	-1PN	$\beta_{\rm ST}$ [59, 151]
Anomalous Acceleration	Extra Dimension Mass Leakage	RS-II Braneworld [152, 153]		-4PN	$\beta_{ m ED}$ [141]
Anomaious Acceleration	Time-Variation of $G$	Phenomenological [137, 154]	-13	-4PN	$eta_{\dot{G}}~[137]$
Scalar Quadrupolar Radiation	Scalar Dipole Field Activation				
Scalar Dipole Force	due to	dCS [140, 155]	-1	+2PN	$\beta_{ m dCS}$ [146]
Quadrupole Moment Deformation	Gravitational Parity Violation				
Scalar/Vector Dipolar Radiation	Vector Field Activation	EA [109, 110], Khronometric [111, 112]		-1PN	$e^{(-1)}$ [112]
Modified Quadrupolar Radiation	due to			0PN	$egin{array}{c} eta_{E}^{(-1)} & [113] \ eta_{E}^{(0)} & [113] \end{array} \end{array}$
Modified Quadrupolar Hadiation	Lorentz Violation		-5	01 IN	$\rho_{E}$ [115]
		Massive Gravity [156–159]	-3	+1PN	
		Double Special Relativity [160–163]	+6	+5.5PN	
		Extra Dim. [164], Horava-Lifshitz [165–167],	+9	+7PN	
Modified Dispersion Relation	GW Propagation/Kinematics	gravitational SME $(d = 4)$ [179]	+3	+4PN	$eta_{ ext{MDR}}$
		gravitational SME $(d = 5)$ [179]	+6	+5.5PN	[145,  156]
		gravitational SME $(d = 6)$ [179]	+9	+7PN	
		Multifractional Spacetime [168–170]	3–6	4-5.5PN	

The Future

**Open** Questions

Multi-Band

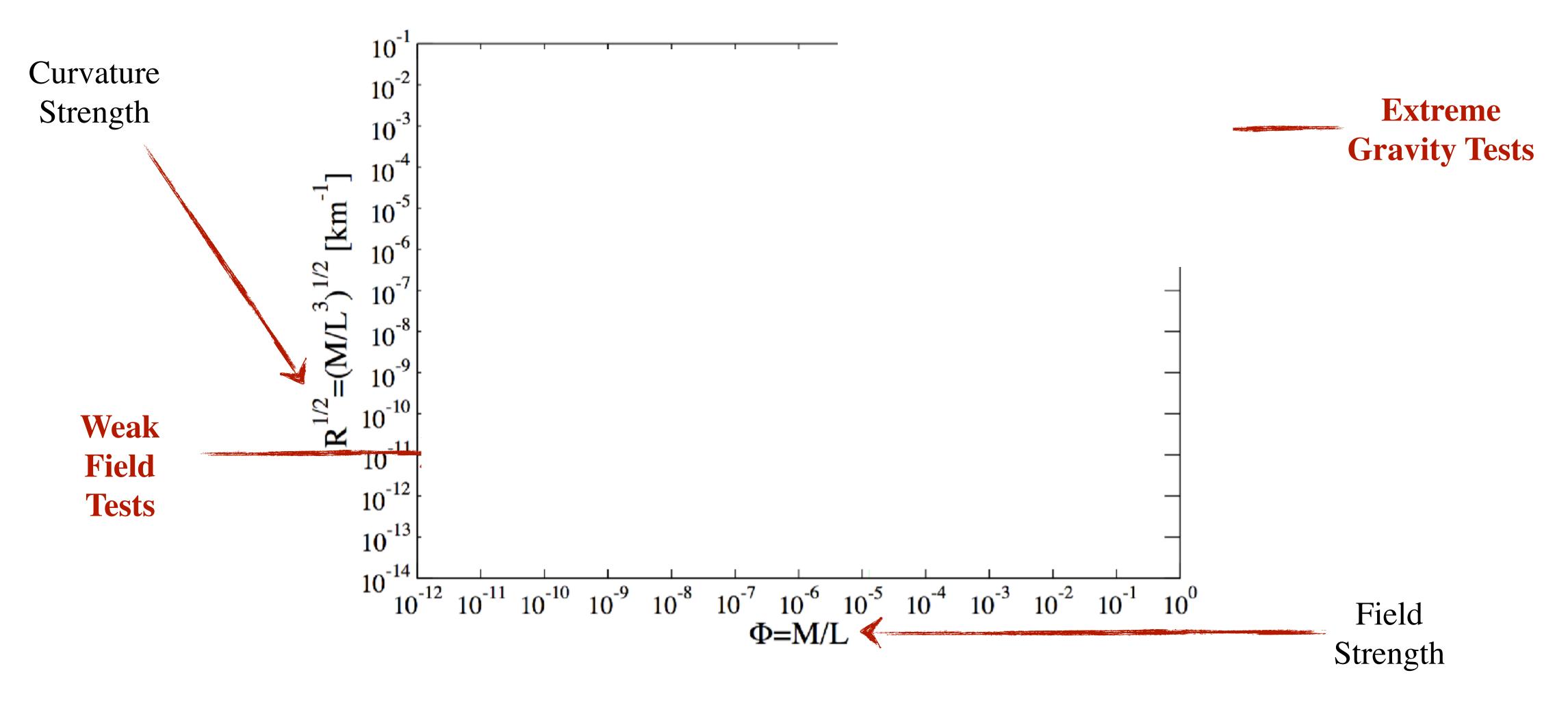
[Yunes & Pretorius, PRD 2009]

[<u>MSU</u>: Cornish et al PRD 84 ('11), Sampson et al PRD 87 ('13), Sampson, et al PRD 88 ('13), Sampson et al PRD 89 ('14), Nikhef: Del Pozzo et al PRD 83 ('11), Li et al PRD 85 ('12), Agathos et al PRD 89 ('14), Del Pozzo et al CQG ('14).]





## Strong Gravity versus Extreme Gravity



[Will, Liv. Rev., 2005, Psaltis, Liv. Rev., 2008, Baker, et al, Siemens & Yunes, Liv. Rev. 2013, Yunes, et al PRD 2016]

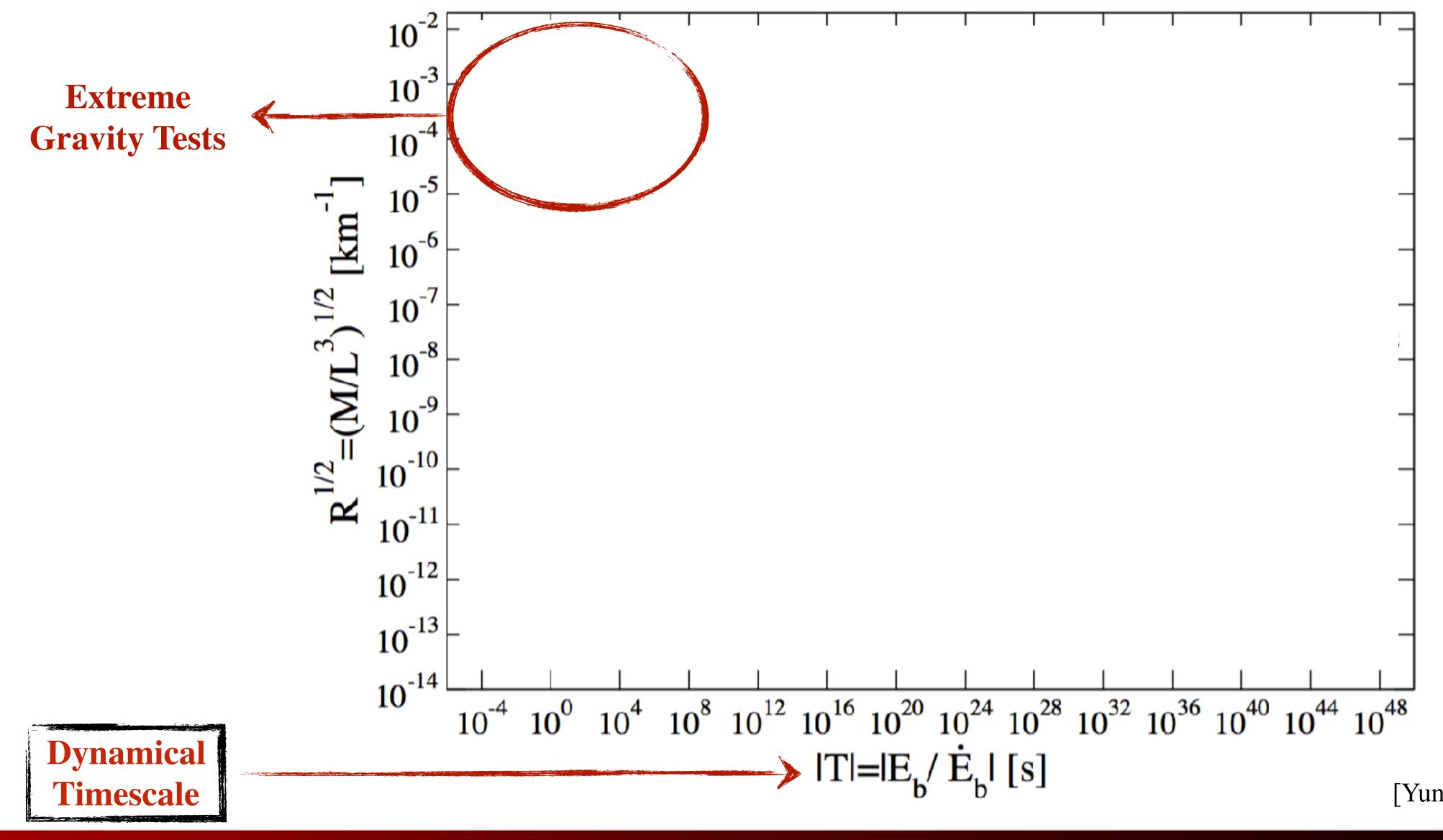








## Extreme Gravity versus Strong Gravity



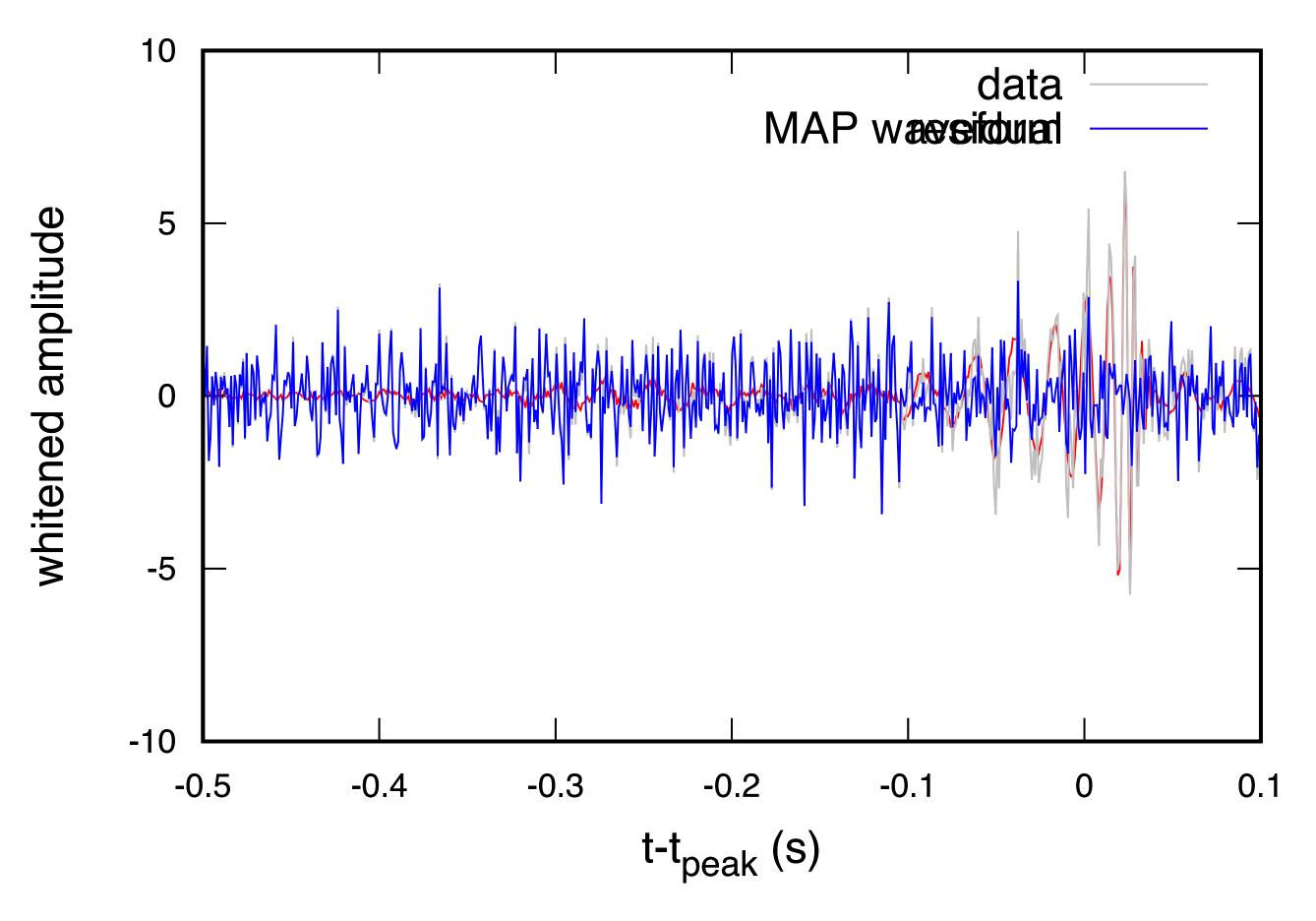
[Yunes, Yagi, Pretorius, PRD '16]







### Residual Consistency



**Open** Questions The Future

Multi-Band

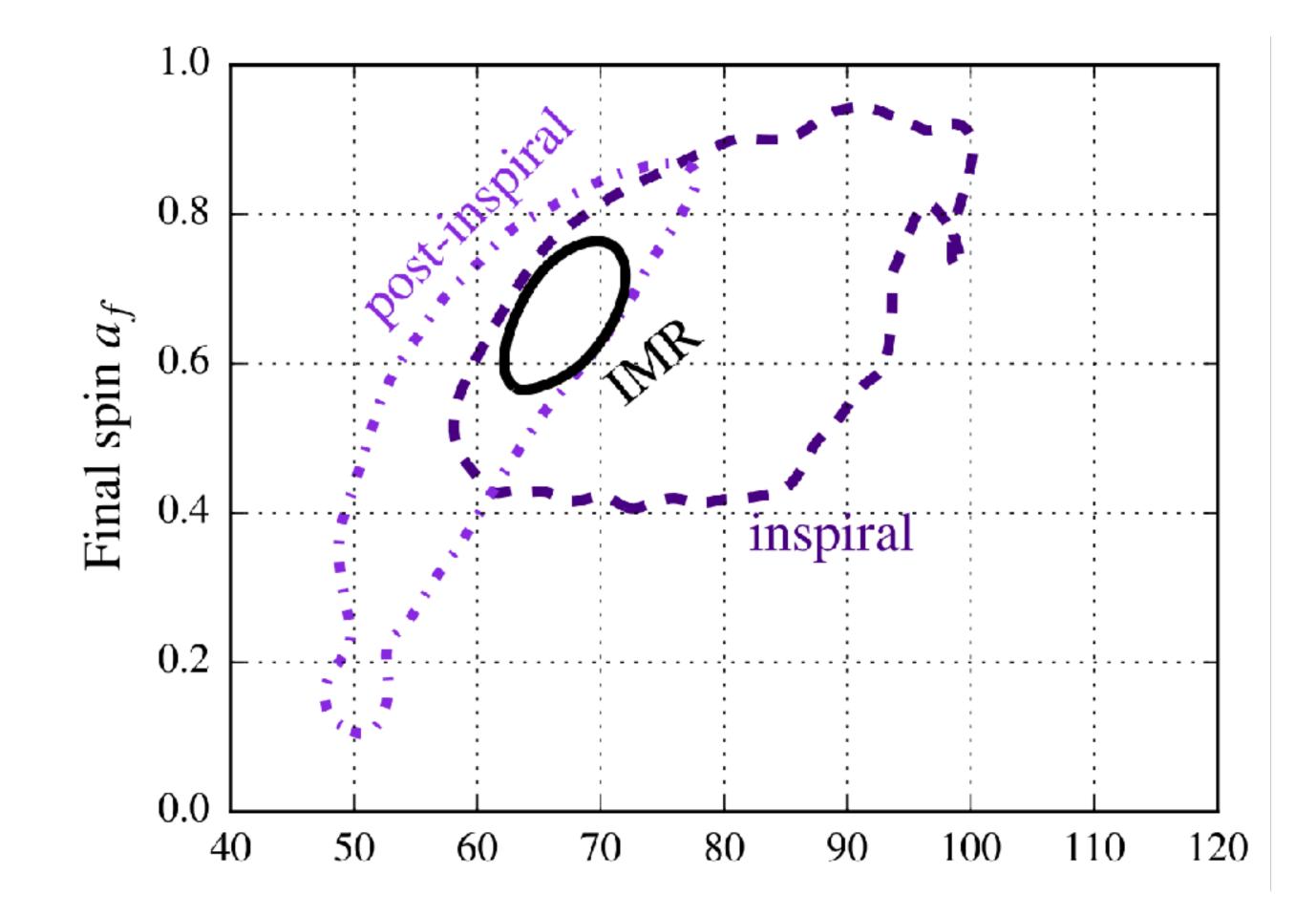
#### [Littenberg & Cornish]

### **SNR of Residual (data - best fit) is consistent with noise**





### (Sort of) Jack-knife Consistency



Multi-Band

#### The Future C

#### Open Questions





### Classifying Deviations

**Gravitational Wave Generation** 

Scalar/Vector Field Activation Gravitational Parity Violation

Gravitational Lorentz Violation

Extra-Dimensional Leakage

Time-Variation of G

Spacetime Dimensionality

Parity Violation

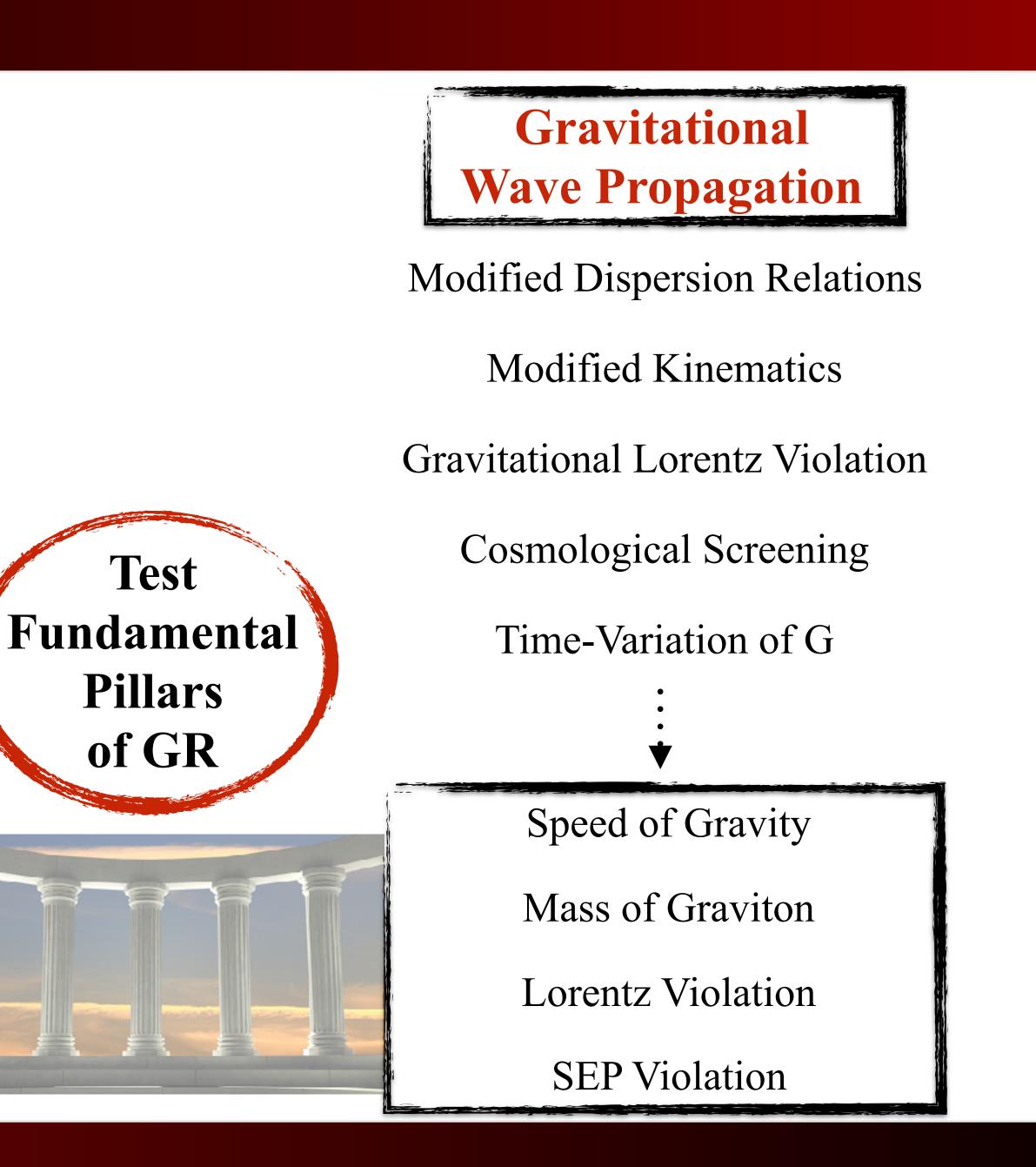
Lorentz Violation

SEP Violation

Multi-Band

The Future

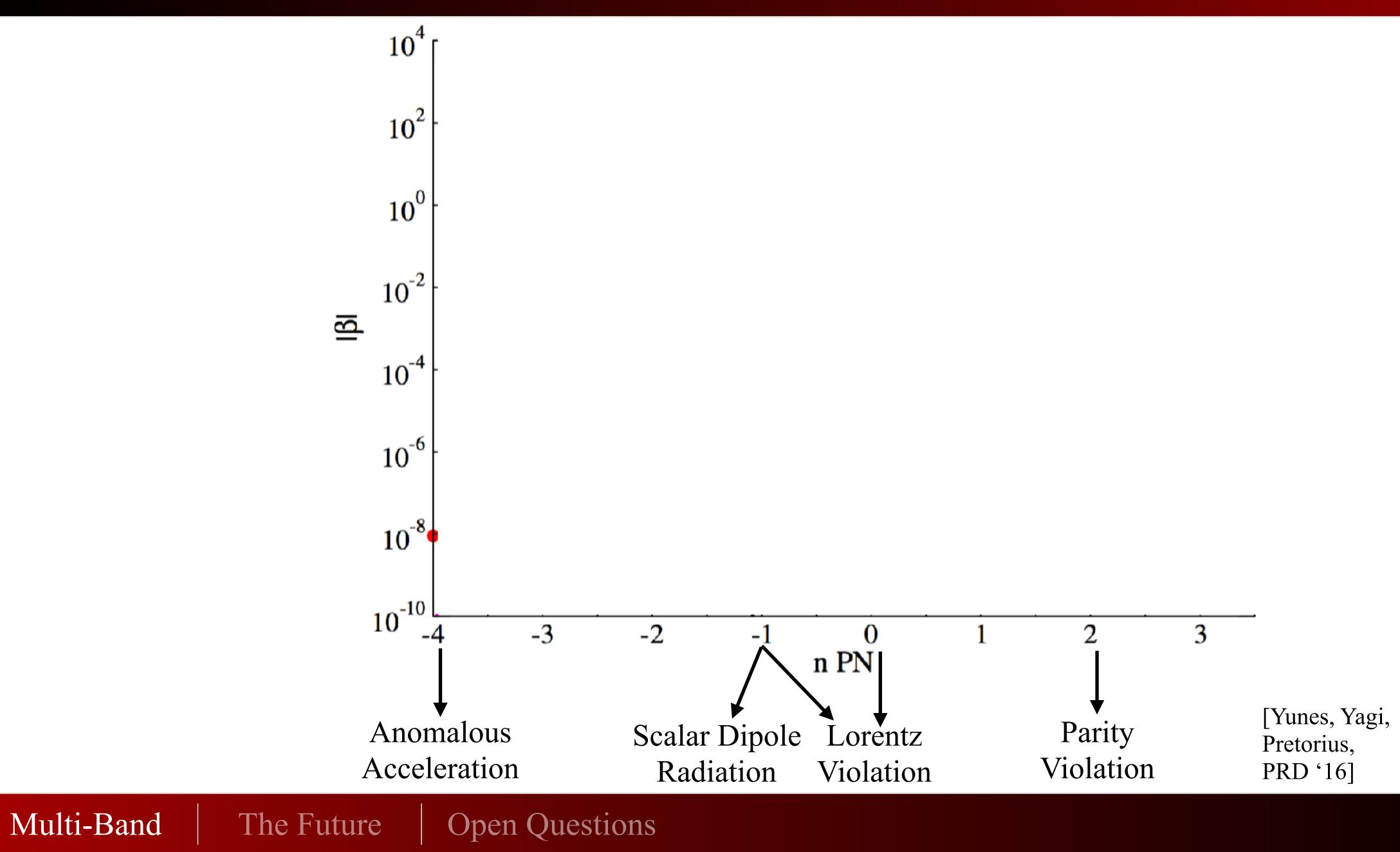
**Open** Questions







### O1 Constraints on GW Generation Effects

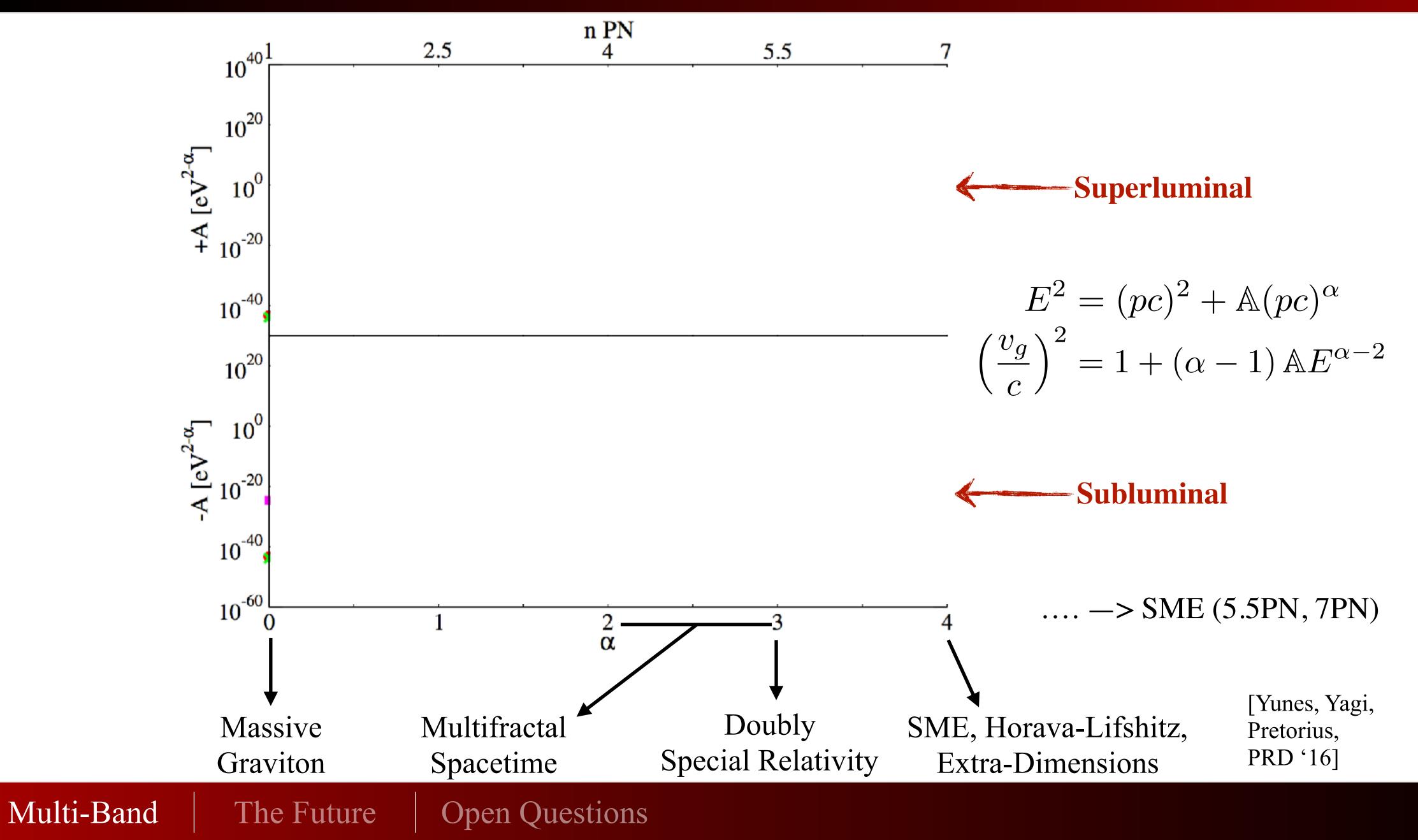








### O1 Constraints on Propagation Effects







## **Theoretical Physics Implications**

Theoretical Mechanism	CR Biller	PN	/	3	Example Theory Constraints							
	Git I mai	11	GW150914	GW151226	Repr. Parameters	GW150914	GW151226	Current Bounds				
Scalar Field Activation	SEP	-1	$1.6 \times 10^{-4}$ $4.4 \times 10^{-5}$	$1.4 \times 10^{-5}$	$\sqrt{ \alpha_{\rm EdGB} }$ [km]		—	$10^7$ [56], 2 [57–59]				
Scalar Field Activation		-1		$4.4 \times 10^{-1}$	$ \dot{\phi} $ [1/sec]			$10^{-6}$ [60]				
Scalar Field Activation	SEP, PI	+2	$1.3  imes 10^1$	4.1	$\sqrt{ lpha_{ m dCS} }$ [km]	—		$10^8 \ [61, \ 62]$				
Vector Field Activation	SED LI	SEP, LI 0	0	0	0	0	$7.2  imes 10^{-3}$	$2.4 \times 10^{-3}$	$(c_{+}, c_{-})$	(0.9, 2.1)	(0.8, 1.1)	(0.03, 0.003) [63, 64]
vector Field Activation	511, 11	0	1.2 × 10	5.4 × 10	$(eta_{ m KG},\lambda_{ m KG})$	(0.42, -)	(0.40, -)	(0.005, 0.1) [63, 64]				
Extra Dimensions	4D	-4	$9.1 imes10^{-9}$	$9.1\times10^{-11}$	$\ell \; [\mu { m m}]$	$\mathbf{5.4  imes 10^{10}}$	$2.0 imes10^9$	$10 - 10^3 \ [65 - 69]$				
Time-Varying $G$	SEP	-4	$9.1  imes 10^{-9}$	$9.1\times10^{-11}$	$ \dot{G} ~[10^{-12}/{ m yr}]$	$5.4 imes10^{18}$	$1.7 imes10^{17}$	0.1-1 [70-74]				

#### **Open** Questions Multi-Band The Future

#### [Yunes, Yagi, Pretorius, PRD '16]







