

# FRIENDLY BLACK HOLES AT THE MUON COLLIDER

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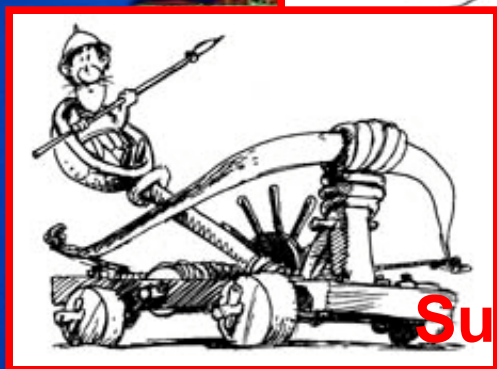
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"The last I heard, Medwick was  
working on a model black hole  
in his lab."

# Hierarchy problem

Observed Planck scale:  $M_{Pl} \sim 10^{19}$  GeV



Susy? Running couplings?

Observed EW scale:  $M_{EW} \sim 10^2$  GeV



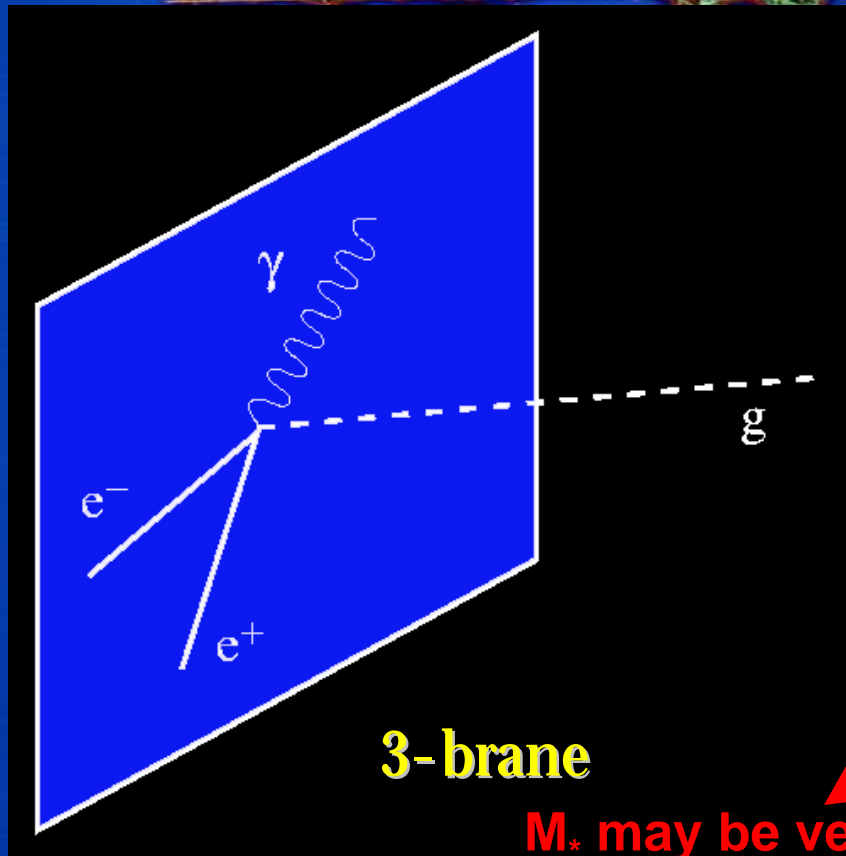
# The approach "from below"

(Arkani-Hamed, Dimopoulos & Dvali 1998; 1999; Antoniadis, Arkani-Hamed & Dvali 1998)

- ◆ The **space time is  $D$ -dimensional ( $D > 4$ )**
- ◆ **SM fields are confined on a 3-brane in a higher-dimensional space time**
- ◆ Only **gravity propagates in the  $n = D - 4$  extra dimensions**



# Brane Universe



$$S = \frac{M_*^{n+2}}{16p} \int d^{n+4}x \sqrt{-g} R(g)$$

$$G_{n+4} = G_4 V_n$$

$$M_{Pl}^2 = M_*^{n+2} V_n$$

**$M_*$  may be very small!!**



# How do we probe extra-dimensions?



# Gravitational effects at the TeV scale

## Perturbative effects (Energy $\lesssim M_*$ )

(Giudice et al. 1998; Mirabelli et al. 1998; Han et al. 1998; Hewett 1999)

- Kaluza-Klein modes
- Virtual graviton exchange

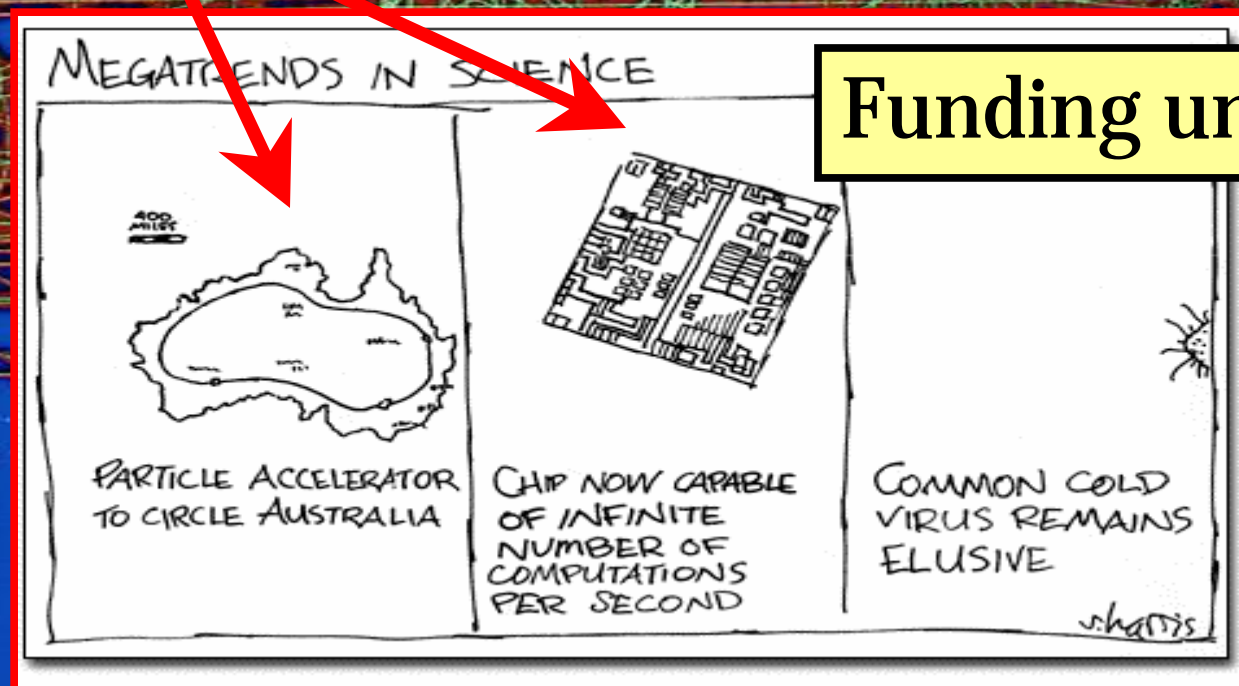
## Nonperturbative effects (Energy $\gtrsim M_*$ )

(Banks & Fischler 1999; Amati et al. 1987)

All depends on  $M_*$ ...

If  $M_*$  is  $\sim 10^{16}$  TeV...

...experimental quantum gravity requires:



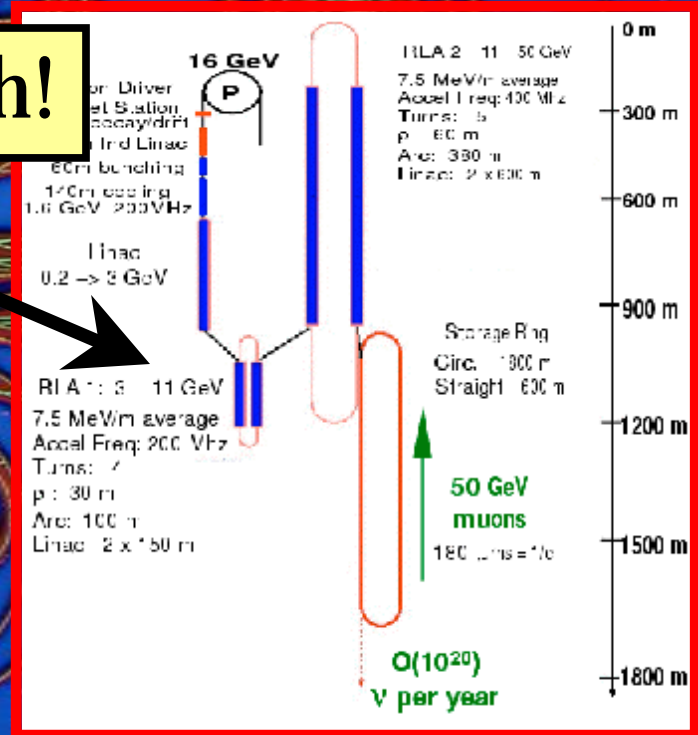
Funding unlikely!

But

If  $M_*$  is  $\sim$  TeV...

...quantum gravity becomes strong at  $\sim$  TeV

**This may be enough!**





# Experimental constraints



- ◆ **Cavendish experiments:  $n=2$ ,  $M_* > 1.6$  TeV**

(Adelberger et al. 2002)

- ◆ **Particle collider experiments:  $M_*/\text{TeV}$  larger than**

(see e.g. Giudice & Strumia 2002; Peskin 2000)

	$n=2$	$n=4$	$n=6$
LEP II	0.90	0.33	0.18
Tevatron	0.86	0.39	0.27
LHC	9.4	3.4	2.1

- ◆ **Astrophysics and cosmology:  $M_*/\text{TeV}$  larger than**

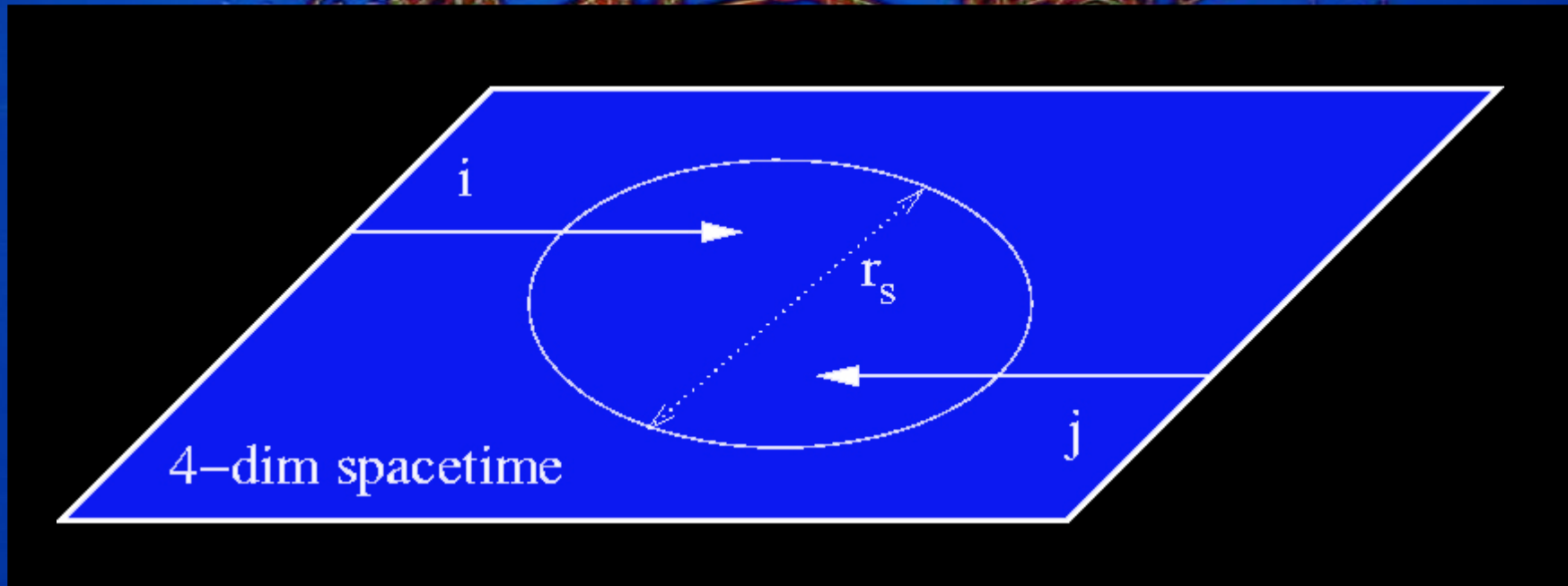
(see e.g. Cullen & Perelstein 1999)

	$n=2$	$n=3$	$n=4$
SN1987A	38-63	2.2-3.9	0.45
Neutron stars	1260	33	
CMBR	65-750	4-32	0.7-4
$C\gamma$ BR	83-263	2.8-7.6	
UHECR		0.2-0.3	0.2-0.3

# Nonperturbative effects:



Scattering of two particles with C-o-M energy  $> \text{TeV}$  and impact parameter  $b < r_s$



## Black hole forms!

# Cross section

(Landsberg & Dimopoulos 2001; Giddings & Thomas 2002)



$$S_{bh}(s; n) \approx pr_s^2 \approx \frac{1}{s_*} \left[ \frac{8\Gamma((n+3)/2)}{(2+n)} \right]^{\frac{2}{n+1}} \left( \frac{s}{s_*} \right)^{\frac{1}{n+1}}$$

where

$$M = \sqrt{s}$$

$$M_* = \sqrt{s_*}$$

- ◆ Uncharged, non-rotating, spherically symmetric BH
- ◆ Cross section = **black disk** (semiclassical regime)
- ◆ Form factor = 1

If  $M \sim \text{TeV}$   $\bar{P} r_s \sim \text{TeV}^{-1}$



- UHECR (neutrino-proton to BH):

$$S_{np \rightarrow bh}(s; n) \approx \sum_i \int_0^1 dx f_i(x, Q) S_{bh}(xs; n)$$

- LHC (proton-proton to BH):

$$S_{pp \rightarrow bh}(s; n) \approx \sum_{ij} \int_0^1 dx \int_x^1 \frac{dy}{y} f_i(y, Q) f_j(x/y, Q) S_{bh}(xs; n)$$

- Muon collider (muon-muon to BH):

$$S_{\mu\mu \rightarrow bh}(s; n) \approx S_{bh}(s; n)$$

# Experimental signatures:



- **Very large and steep cross section**

(Landsberg & Dimopoulos 2001; Giddings & Thomas 2002)

- ◆ **Visible transverse energy / High sphericity events**

(Giddings & Thomas 2002)

- ◆ **High (?) multiplicity events (hadronic jets + leptons + hard quanta at the end of decay)**

(Giddings & Thomas 2002; Cavaglia` 2003)

- ◆ **Ratio of hadronic to leptonic activity ~ 5:1**

(Giddings & Thomas 2002; Han et al. 2002; Cavaglia` 2003)

- ◆ **Possible large missing energy**

(Cavaglia`, Das & Maartens 2003)

- ◆ **Suppression of hard perturbative scattering processes**

# Example

$M_* = 1 \text{ TeV}, n = 6, M_{\text{BH}} = 12 \text{ TeV}$

quark	5	gluon	1
charged $l$ .	1	W	0
neutrino	0	Z	0
Higgs	0	graviton	0
photon	0		

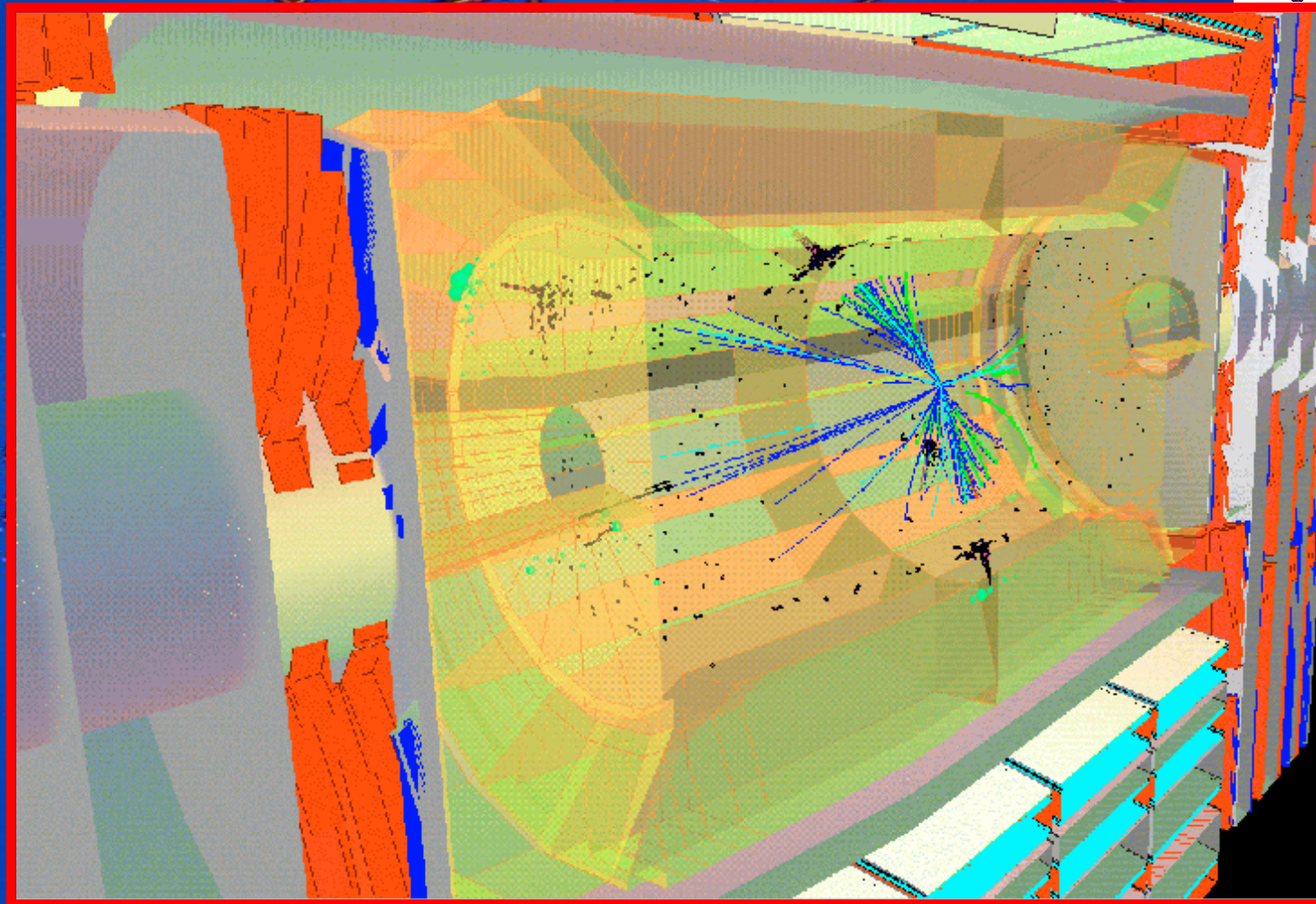
Quarks and gluon hadronize. H-to-L ratio  $\sim 5:1$

# Cross section uncertainties



- ◆ **Nonrelativistic limit estimates**
- ◆ **Classical photon capture / plunging estimates**  
(Berti, Cavaglia` & Guattieri 2003)
- ◆ **Collisional energy loss/inelasticity**  
(Yoshino & Nambu, 2003)
- ◆ **Angular momentum correction**  
(Yoshino & Nambu 2003)
- ◆ **Charge effects**  
(Casadio & Harms 2002)
- ◆ **Minimal BH formation mass**  
(Cavaglia`, Das & Maartens 2003)

# BH event at CMS



(Model of Landsberg & Dimopoulos, Herwig implementation by C. Harris & P. Richardson, generated by A. de Roeck, simulated & visualised by S. Wynhoff)



# The LHC vs the Muon Collider

LHC: Center-of-mass energy (pp) = 14 TeV

(PbPb) = 5.5 TeV

mC: Center-of-Mass energy = 4 TeV

**Disadvantage!**  
VLHC: Center-of-Mass energy = up to 100 TeV

VLmC: Center-of-Mass energy = up to 200 TeV

Cross section at the LHC:

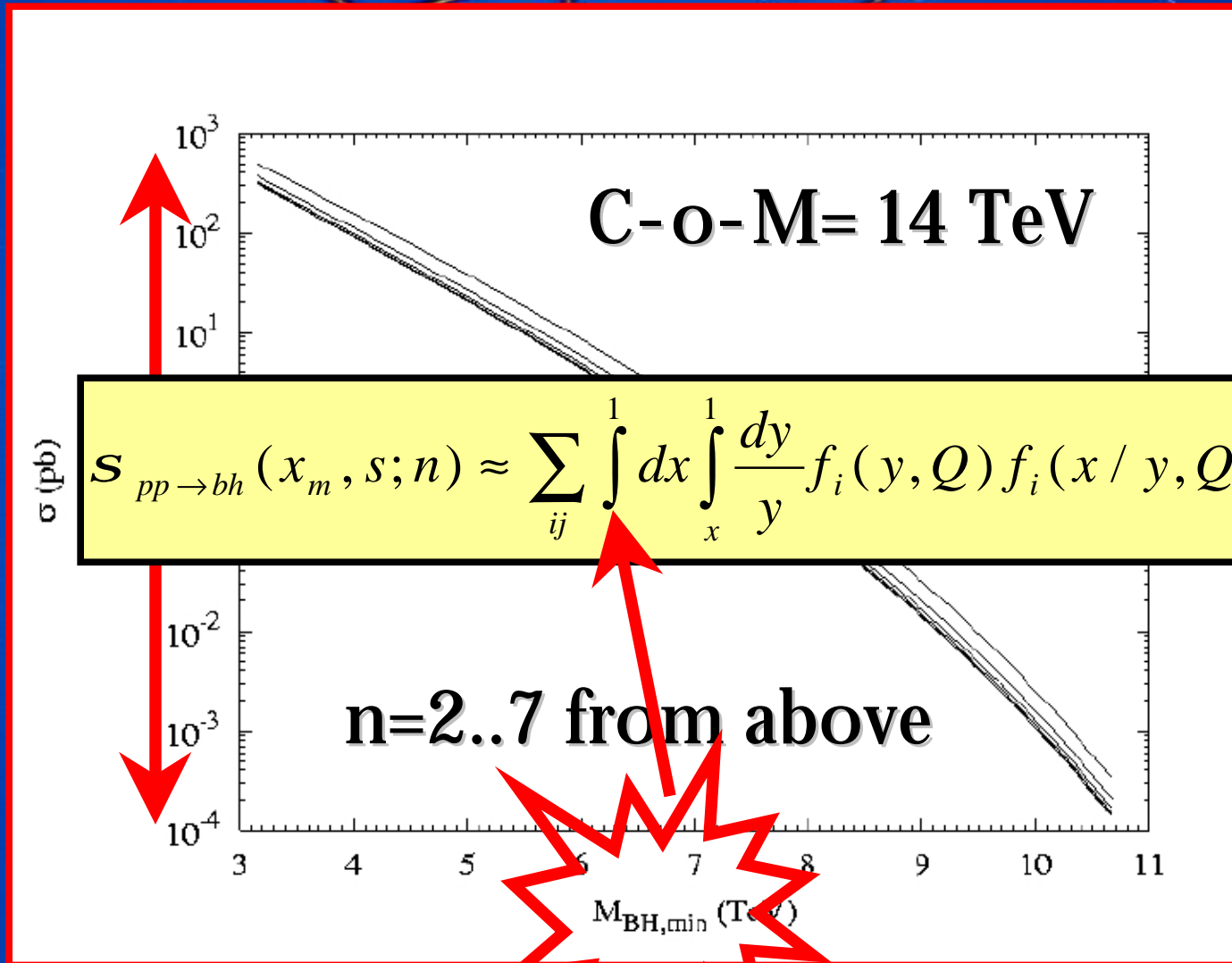
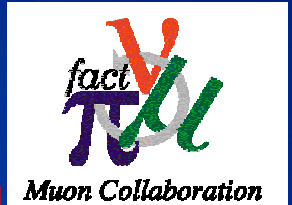
Up to  $10^4$  pb

Cross section at the 4-TeV Muon  
collider:

**Disadvantage!**

$\sigma \approx 10^2$  pb

# BH cross section at the LHC:

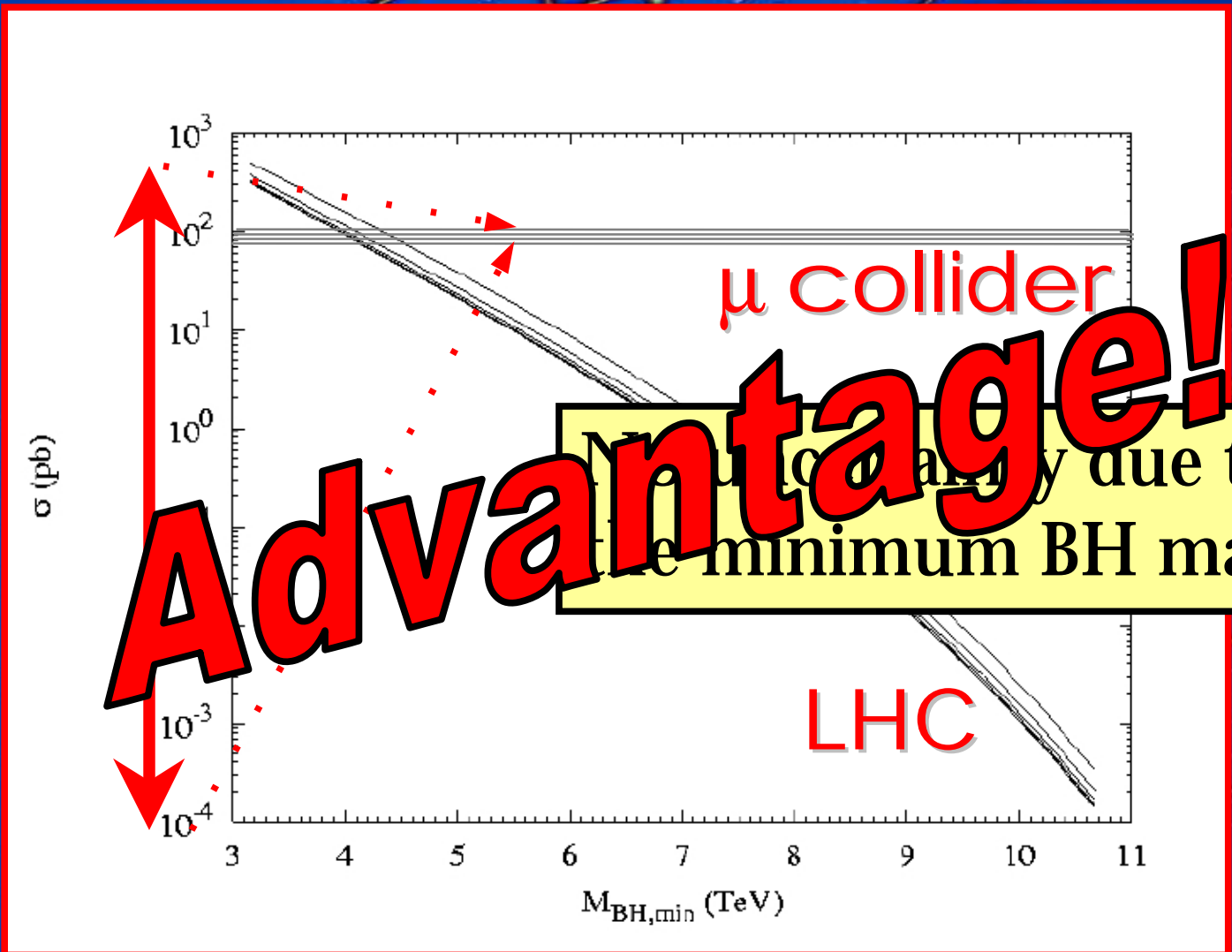


# BH cross section at the muon collider:

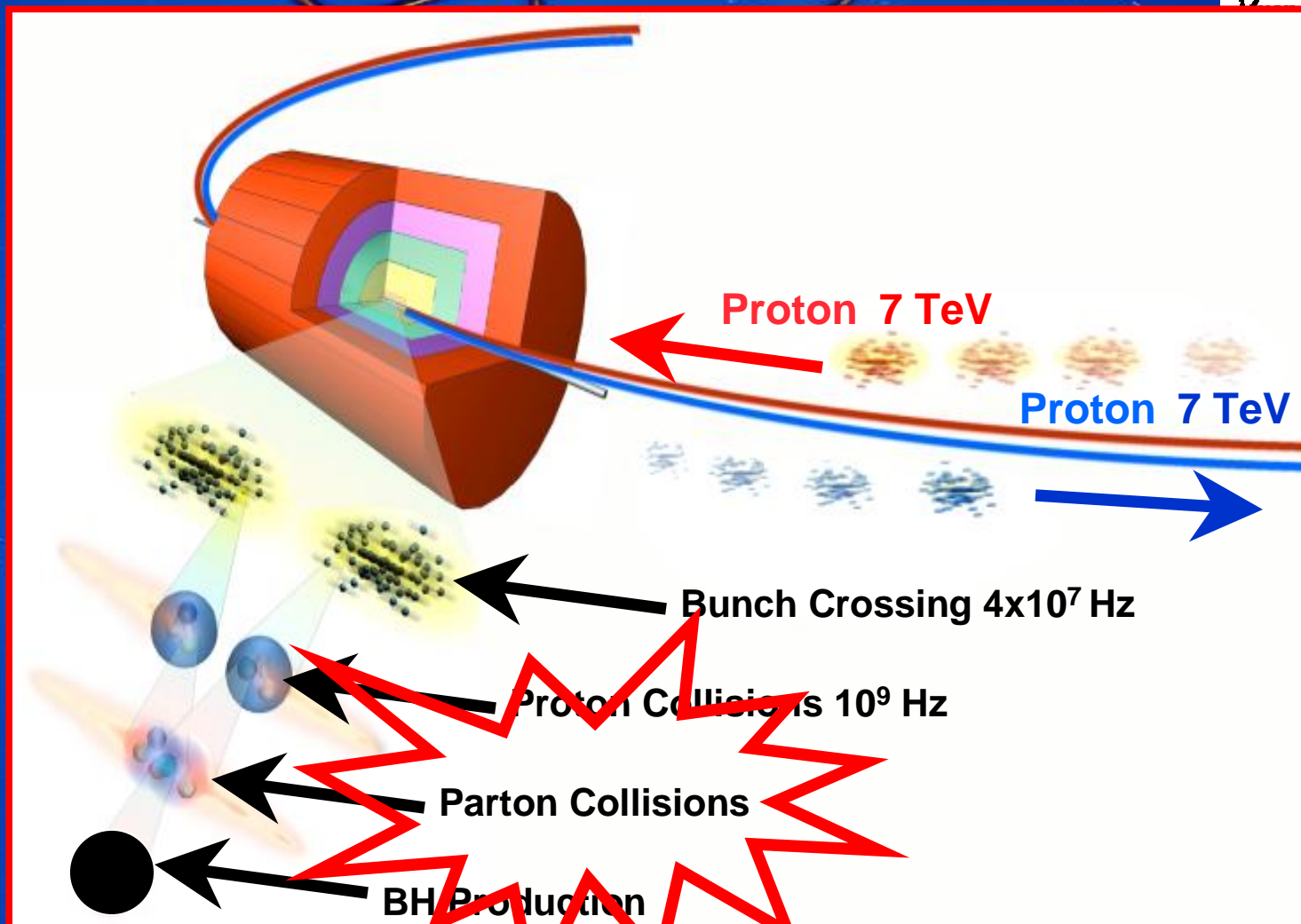
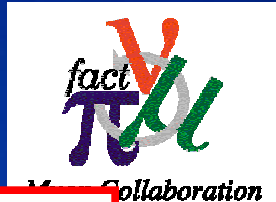
$$S_{mm \rightarrow bh}(s; n) \approx \int d^2e f(e_1) f(e_2) S_{bh}[s(e_1, e_2); n]$$

$$f(e_1) \approx d(e_1 - e_0), \quad f(e_2) \approx d(e_2 - e_0)$$

If  $\varepsilon_m \ll \varepsilon_0$ , the cross section does not depend on the minimum BH mass



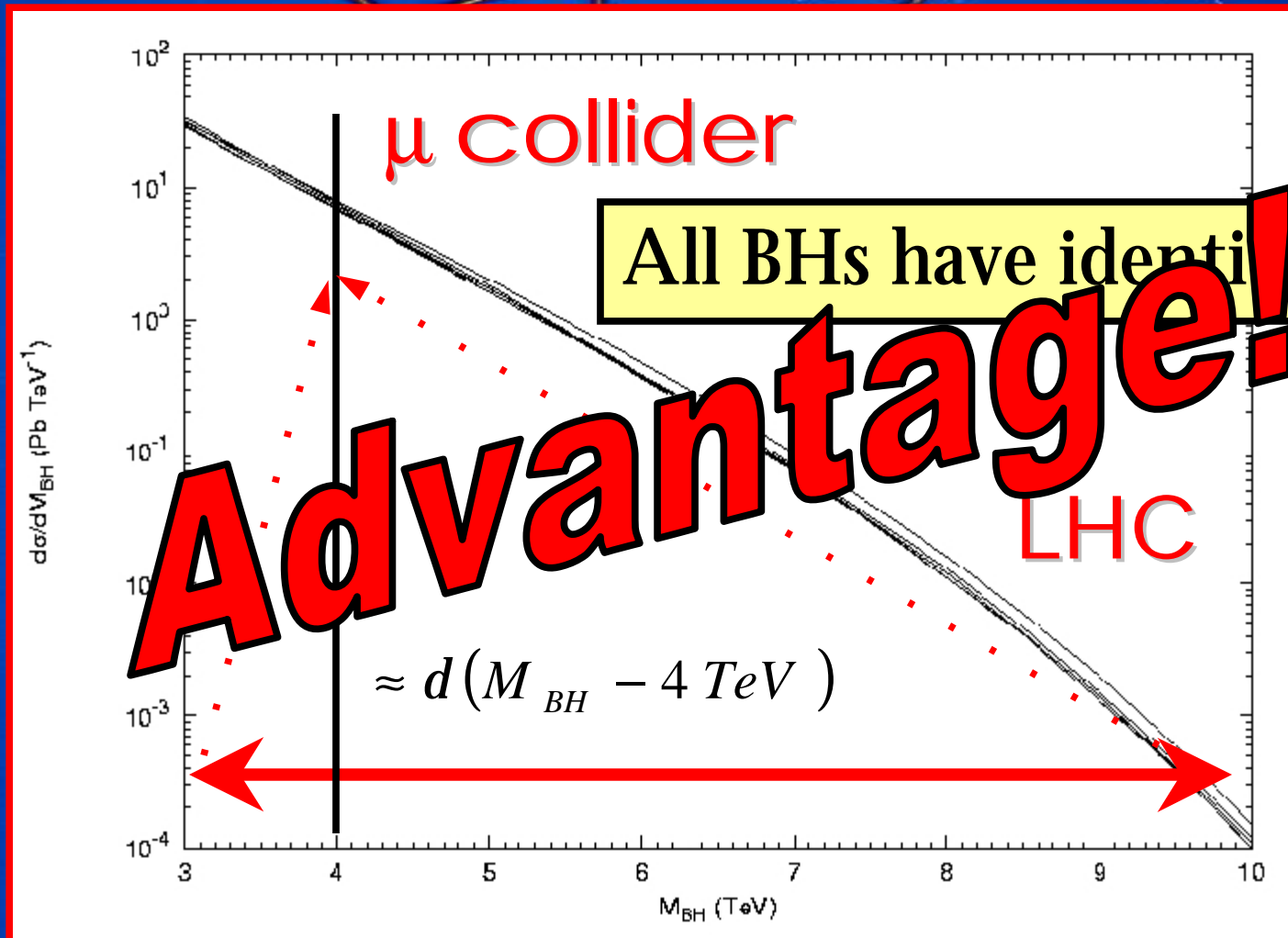
# BH collisions at LHC



# Differential cross section at the LHC

$$\frac{d\mathcal{S}_{pp \rightarrow bh}(x, s; n)}{dM_{BH}} \approx \int_x^1 \frac{dy}{y} f_i(y, Q) f_i(x/y, Q) \mathcal{S}_{bh}(xs; n)$$

# Distribution of BH masses





# What we (want to) do:

- ◆ **Compute the cross section** taking into account the beam profile
- ◆ **Systematical analysis** of the parameter space
- ◆ **Write an improved montecarlo for BH events** at the LHC and at the muon collider
- ◆ **Simulate BH events** at the LHC and at the muon collider
- ◆ **Analysis and comparison**



# Conclusion



- ◆ If large extra dimensions exist  $\Rightarrow$  Planck scale  $\sim$  TeV
- ◆ Nonperturbative quantum gravity effects at  $E \sim$  TeV
- ◆ Creation of black holes & branes at the muon collider!
- ◆ If not, constraints on the Planck scale. If yes, new physics at the TeV scale
- ◆ Possible tests of strong gravitational effects:
  - Hawking radiation
  - Generalized uncertainty principle

