

# Micro black hole back-to-back decay signature searches

Octavian Micu

Institute of Space Science, Bucharest, Romania

In collaboration with:

X. Calmet (U. Sussex),

N. Arsene & L. Caramete (ISS)

Seventh Gulf Coast Gravity Meeting

University of Mississippi, Oxford (MS)

April 20, 2013



# TeV scale Black Holes

- In 4D the scale of gravity is approximately  $10^{16}$  TeV.
- TeV range micro Black Holes appear in brane world models.
  - Standard Model fields live on our “brane” while gravity propagates in the large extra-dimensional volume and so the Planck scale can be anywhere between a few TeV and  $10^{16}$  TeV.
  - Black holes can have any mass above  $M_{PL}$ :
    - $M_{PL} < M < 5 M_{PL}$ : quantum black holes;
    - $M > 5 M_{PL}$ : semi-classical black holes.
  - *Hoop conjecture*: “a black hole forms whenever the impact parameter  $b$  of two colliding objects (of negligible spatial extension) is shorter than the radius of the would-be-horizon (roughly, the Schwarzschild radius, if angular momentum can be neglected) corresponding to the total energy  $M$  of the system.”

$$b \lesssim \frac{2 l_{Pl} M}{M_{Pl}}$$

# Non-thermal Black Holes @ the LHC

No non-thermal Black holes @ the LHC!  
 ...so far! (upgrades in progress!)

- At cosmic ray observatories:
  - ✓ Ground based;
  - ✓ Space based.
- Neutrino telescopes.

## Non-thermal Black Holes (QBH)

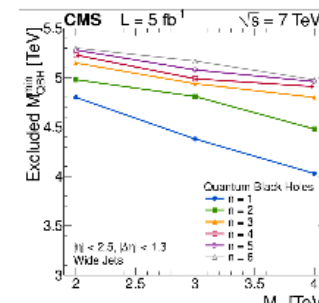
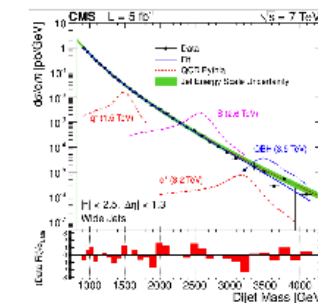
- LHC parton energy needs to be high relative to  $M_D$  for black hole to Hawking evaporate thermally.
- Black holes with threshold mass  $M_{th}$  near  $M_D$  probably do not decay thermally.
- Dijet decays assumed so far.
- QBH appears as bump in dijet mass spectrum.
- Increased discrimination power by using angles.
  - Define central region  $\chi = \exp(|y_1 - y_2|) < 30.0$ .

$$F_\chi(m_{jj}) \equiv \frac{dN_{\text{central}}/dm_{jj}}{dN_{\text{total}}/dm_{jj}}$$

18 March 2013

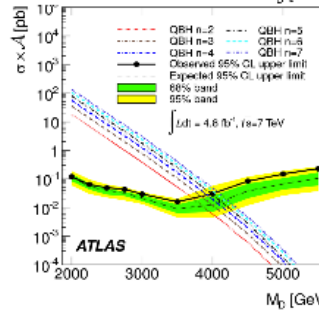
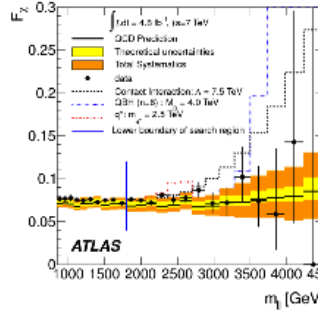
Doug Gingrich (COST Workshop)

## Non-thermal Black Holes Results



JHEP 01 (2013) 013

$M_{th} > 4-5.3$  TeV  
 ( $n = 1-6$ ,  $M_{th} > M_D$ ,  
 $M_D = 2, 3, 4, 5$  TeV)



JHEP 01 (2013) 029

$M_{th} > 3.71-4.07$  TeV  
 ( $n = 2-7$ ,  $M_{th} = M_D$ )

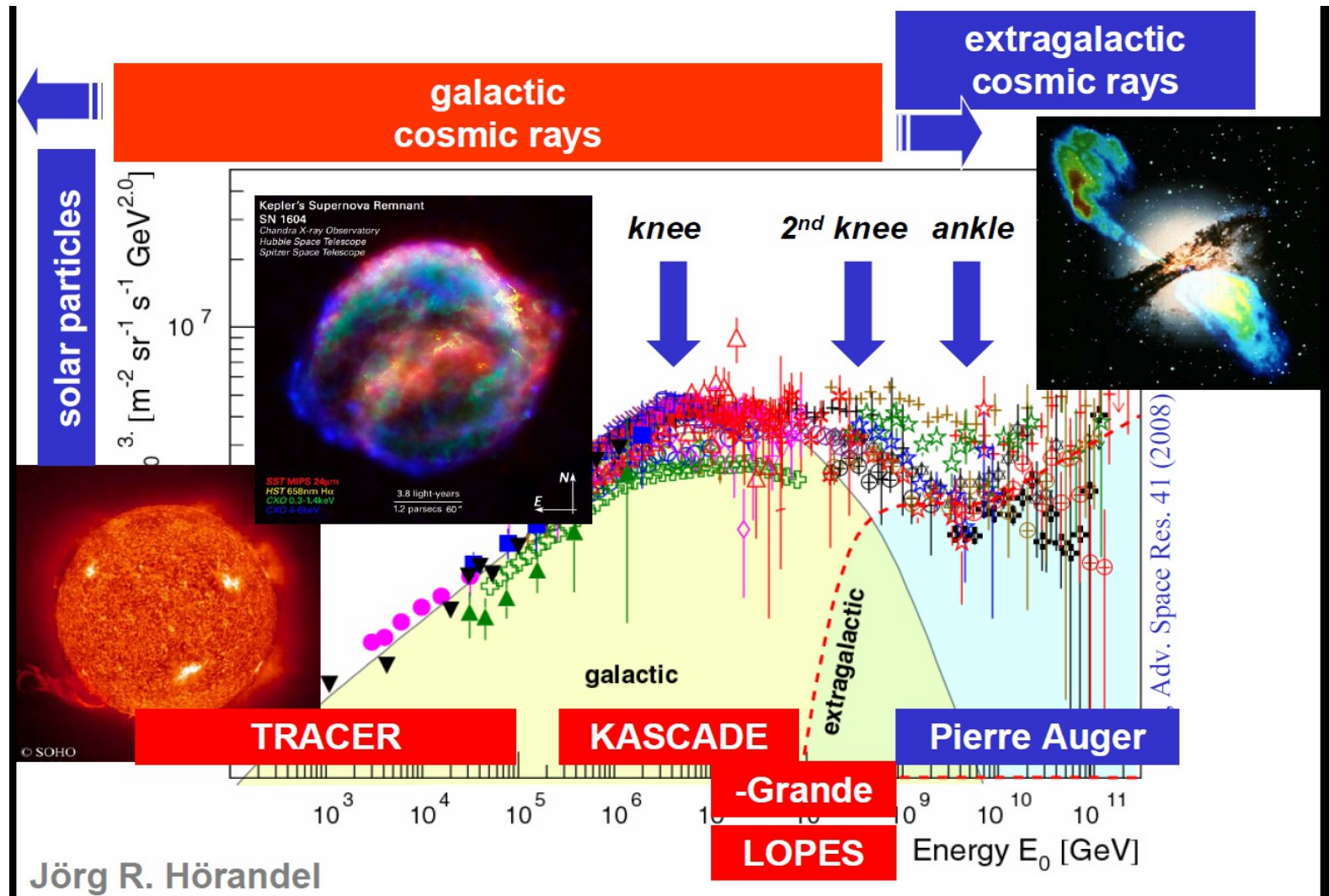
18 March 2013

Doug Gingrich (COST Workshop)

18/21

# Micro black holes in cosmic rays data

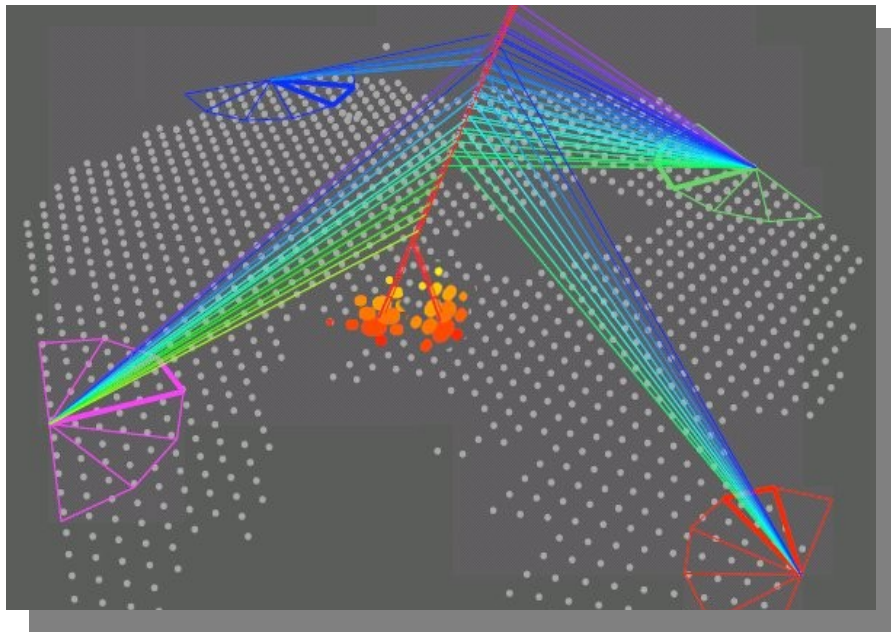
- What other ultra-high energy data do we have available?



# Micro black holes in cosmic rays & neutrino data

"A long time ago in a galaxy far, far away..."

...about a year ago at a "Black Holes in a Violent Universe" COST meeting (European Cooperation in Science and Technology), the idea of a unique micro black hole signature which could be discovered in the cosmic ray data came up.



- If micro black holes with the masses close to the Planck mass, which are non-thermal, are formed in high energy collisions, the main decay channel for these objects is a „back-to-back“ decay into two particles!
- If black holes form as a result of the impact between UHECRs or neutrinos with particles in the upper atmosphere, it is possible for simultaneous double shower events to be observed!

# Micro black holes in cosmic rays & neutrino data

X. Calmet, L.I. Caramete and O. Micu. JHEP **1211**, 104 (2012); arXiv:1204.2520 [hep-ph].

- The number of events expected to be seen by a cosmic rays experiment:

$$N = \int dE N_A \frac{d\Phi}{dE} \sigma(E) A(E) T$$

- Assuming the collision between a particle  $(\gamma_1 m_1 c, \gamma_1 m_1 \vec{v})$  with a particle  $(m_2 c, 0)$  at rest, the resulting black holes...

- Have a mass of:

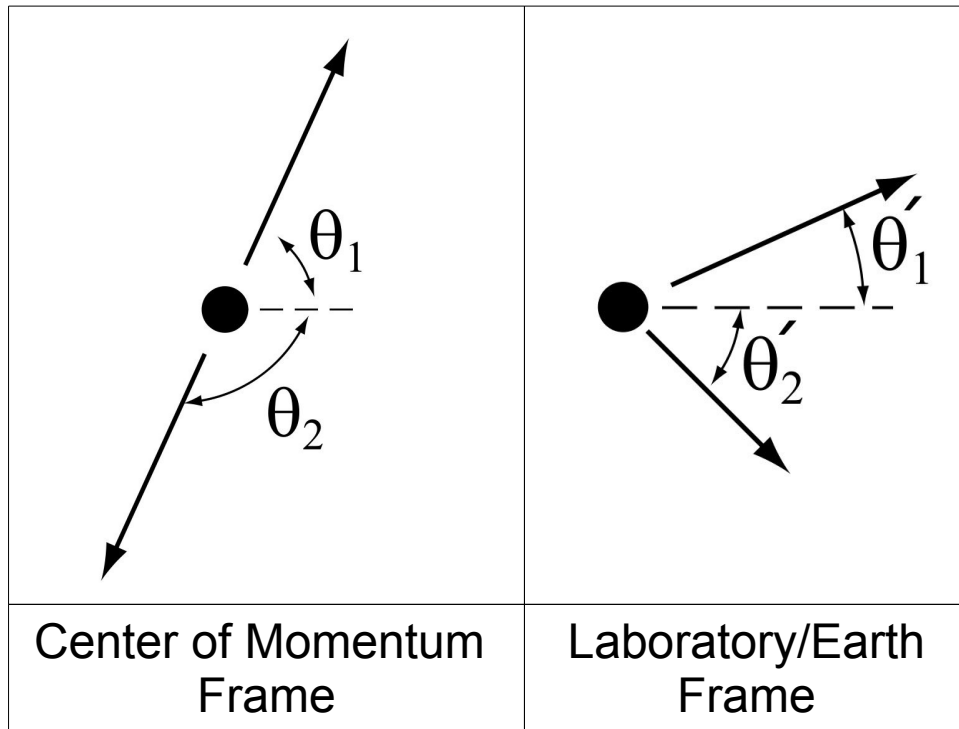
$$M_{BH} = \sqrt{m_1^2 + m_2^2 + 2\gamma_1 m_1 m_2}$$

- And move relativistically with a Lorentz factor of:

$$\gamma_{BH} = \frac{\gamma_1 m_1 + m_2}{M_{BH}}$$

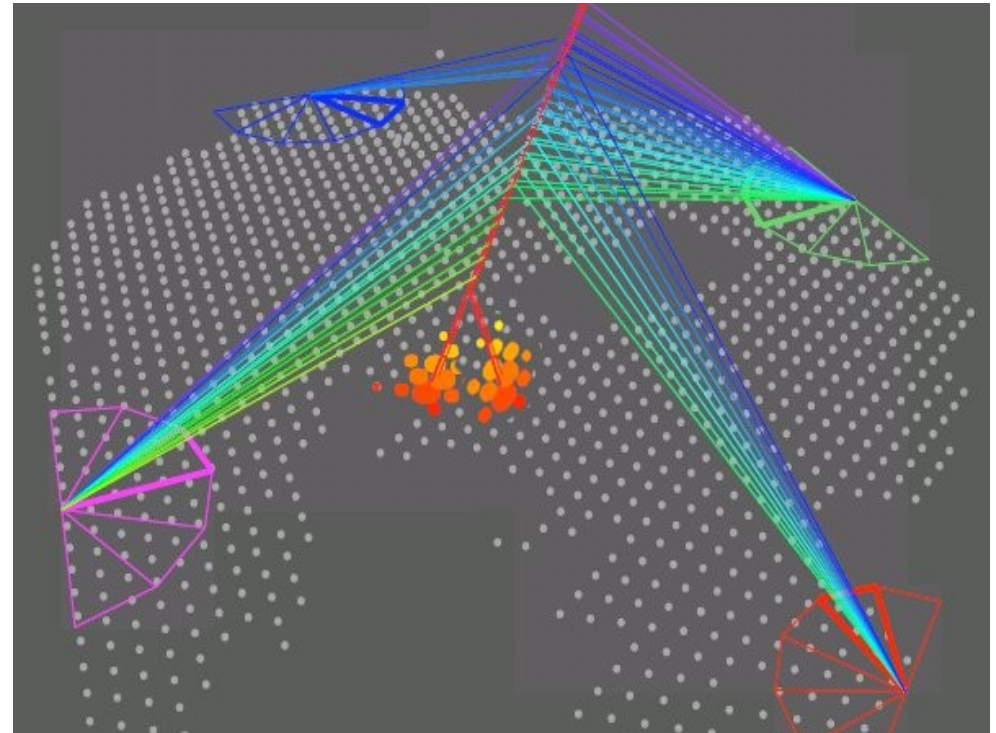
UHECR Energy (TeV)	Black hole mass (TeV)
$10^5$	7
$10^6$	25
$10^7$	79
$10^8$	250

# Micro black holes in cosmic rays & neutrino data



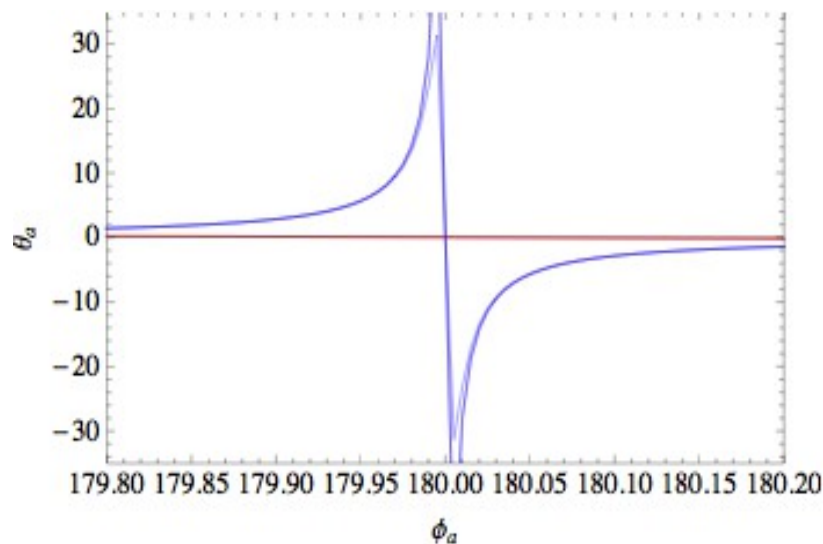
• Lorentz transformed angles

$$\tan \theta'_i = \frac{\sin \theta_i}{\gamma_{BH} \beta_{BH} \frac{E_i}{p_i} + \gamma_{BH} \cos \theta_i}$$



# Micro black holes in cosmic rays & neutrino data

- Case study: Pierre Auger Observatory
  - ✓ For the Planck mass in the 10 TeV range,
  - ✓ UHECR or neutrino energies larger than  $10^6$  TeV
  - ✓ Back-to-back decay into two standard model particles



Angle  $\theta_1$  as function of the mass  $m_1$  for  $m_d = 4.8$  MeV,  $m_{\text{muon}} = 105.7$  MeV,  $m_{\text{tau}} = 1.777$  GeV (all in blue) and  $m_{\text{top}} = 171$  GeV (in red).

- Two showers can be observed for about 0.11% of the total number of back-to-back decays!
- These events are rare but one might still hope to observe them.

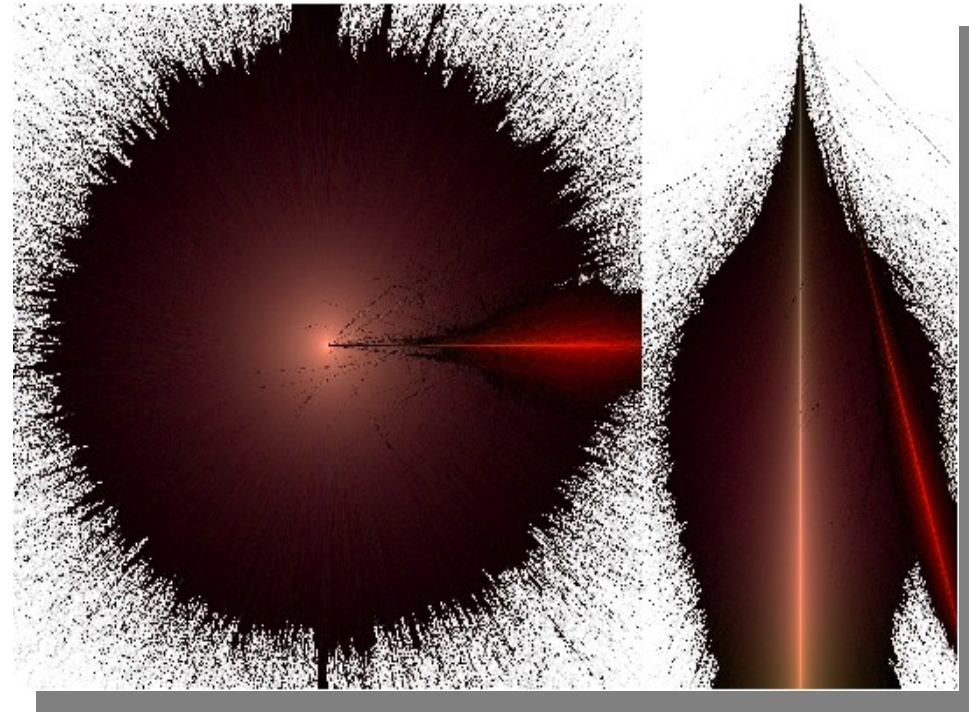
No. of extra dimensions	No. of events/year
1	81
4	460
5	609
6	765
7	925



# Back-to-back decay signature @ UHECR observatories

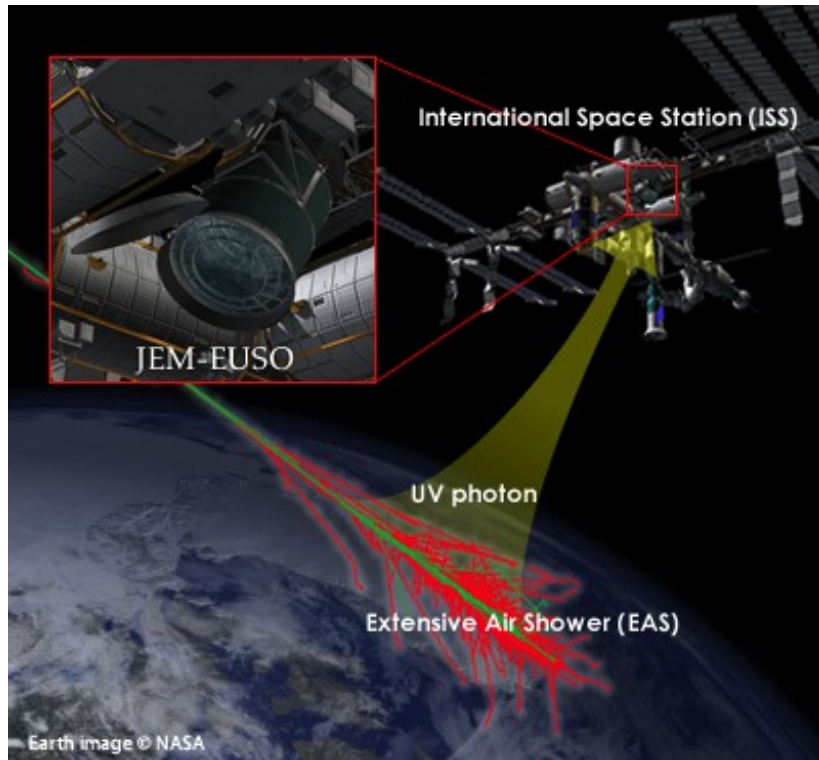
## Back-to-back decay simulations CORSIKA COsmic Ray Simulations for Kascade (20° separation between the showers)

- × Showers overlap for large parts of their development!
- × For smaller angles they overlap even more!
- × The identification of the two separate showers by the ground detectors depends on the altitude at which they are produced!
- × One will have to also tell apart oval shaped imprints due to two showers from single showers oriented at an angle wrt. the Earth.



CORSIKA simulations of two showers ( $10^{13}$  eV &  $9 \times 10^{16}$  eV) developing from a height of 30 Km. The angle between the showers in the Earth frame is 20°.

# Back-to-back decay signature @ space observatories



## **JEM-EUSO**

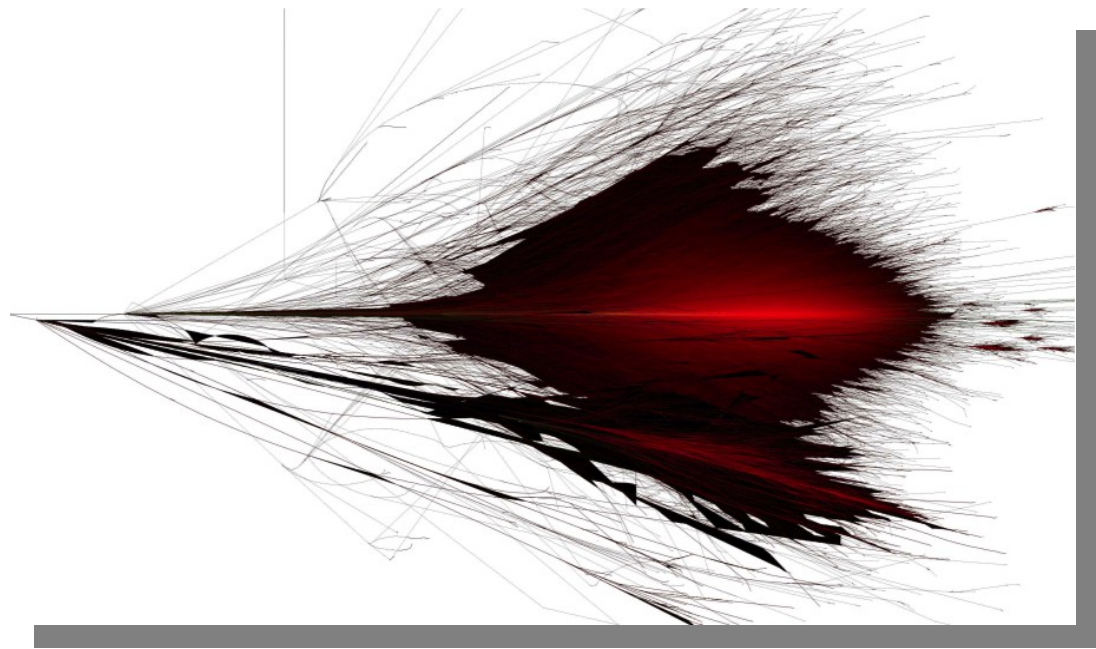
Extreme Universe Space Observatory

<http://jemeuso.riken.jp/en/>

Will observe a volume 50 – 250 times larger than the Pierre Auger Observatory.

Advantages:

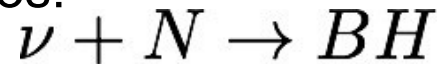
- A more than 20 times larger acceptance!
- The entire fluorescence showers would be visible which would make the features of a two-shower event to be distinguishable!



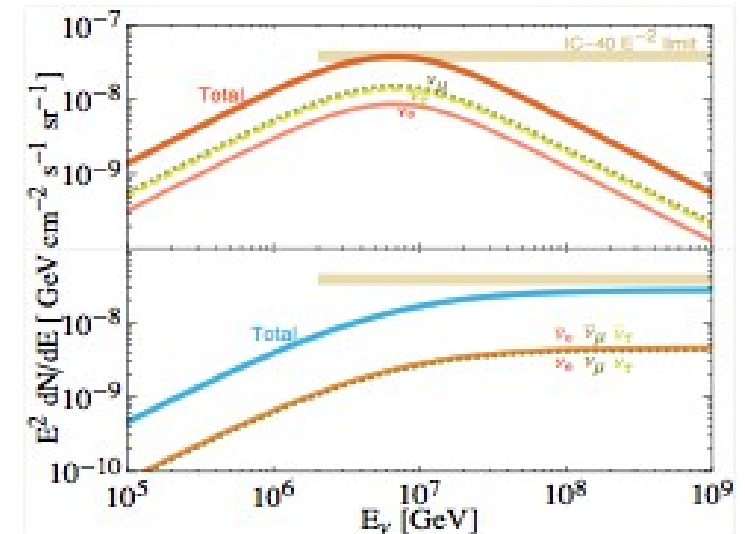
# Back-to-back decay signature @ neutrino observatories

N. Arsene, X. Calmet, L.I. Caramete and O. Micu. arXiv:1303.4603 [hep-ph].

- QBHs produced by highly energetic neutrinos:



- Neutrino observatories (telescopes)
  - ✓ Detect the muons induced by high energy neutrinos (Cherenkov light).
  - ✓ The direction of propagation is derived with high accuracy!
  - ✗ There are many models which lead to very different values for the flux at high energies!!!



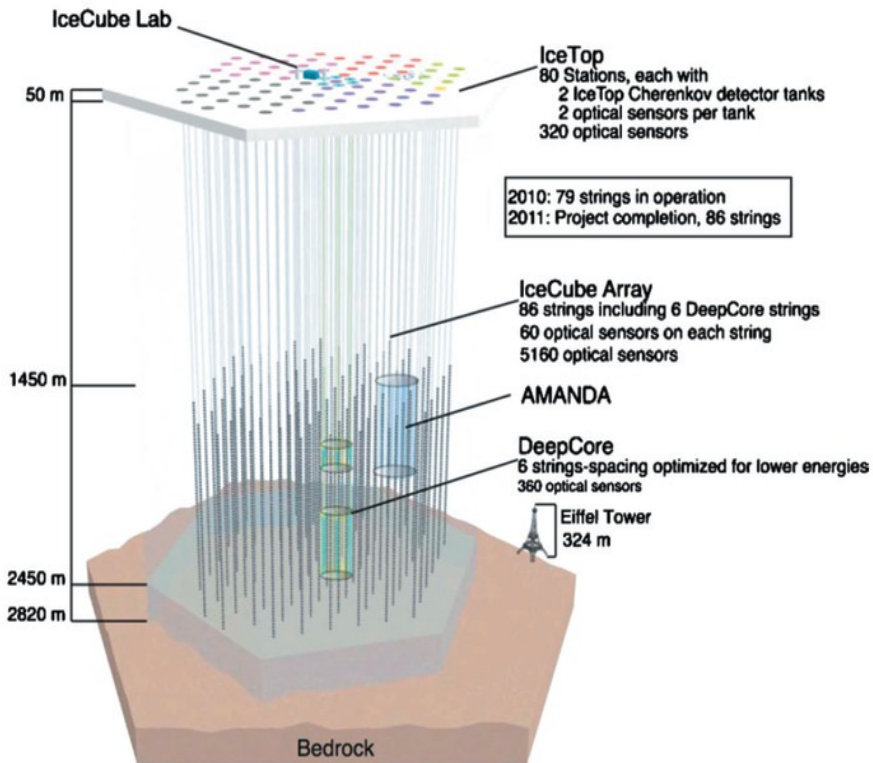
Neutrino flux at high energies:

M. Kistler, T. Stanev, H. Yüksel.

arXiv:1301.1703 [astro-ph]

- The range of interaction lengths for neutrino energies between  $10^{16}$  -  $10^{18}$  eV is between  $6.6 \times 10^3$  km and  $9.4 \times 10^2$  km water equivalent in rock!!!
  - The Earth is opaque for neutrinos with these energies and experiments can see neutrinos coming from above or the ones skimming the Earth!

# Back-to-back decay signature @ neutrino observatories



## IceCube

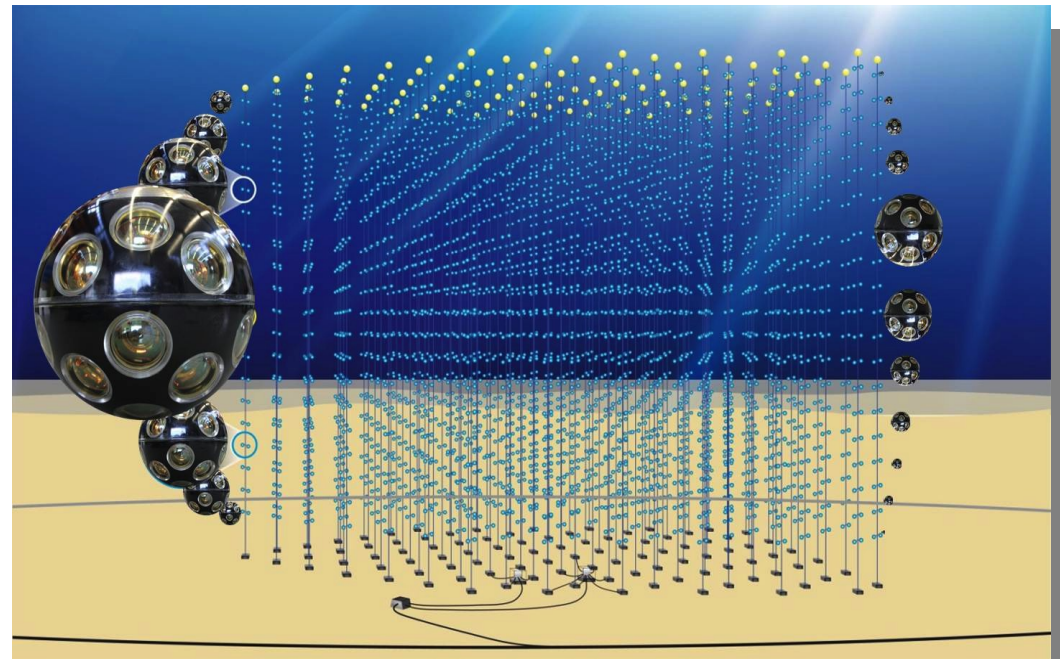
South Pole Neutrino Observatory.

<http://icecube.wisc.edu/>

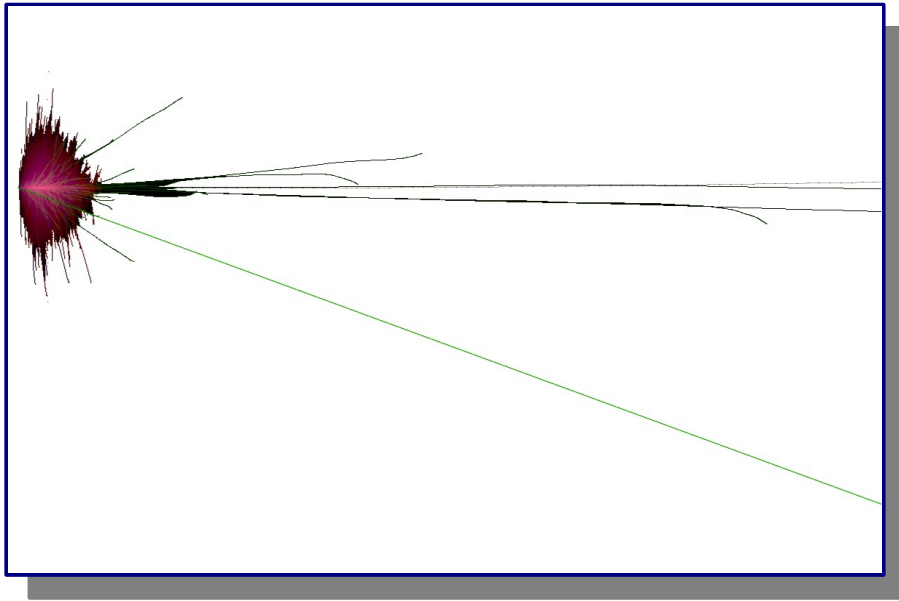
## KM3NeT

<http://www.km3net.org>

...future multi km<sup>3</sup> sized neutrino telescope to be built on the bottom of the Mediterranean Sea.



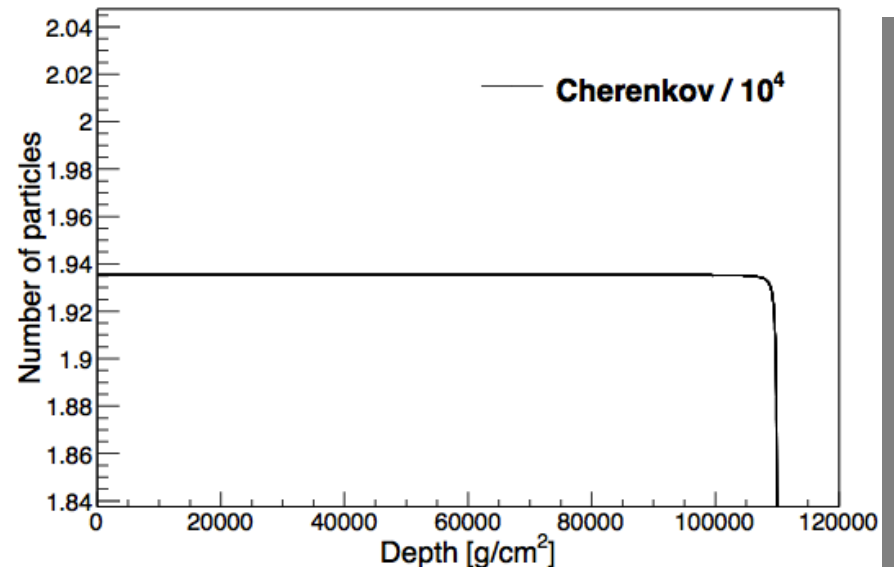
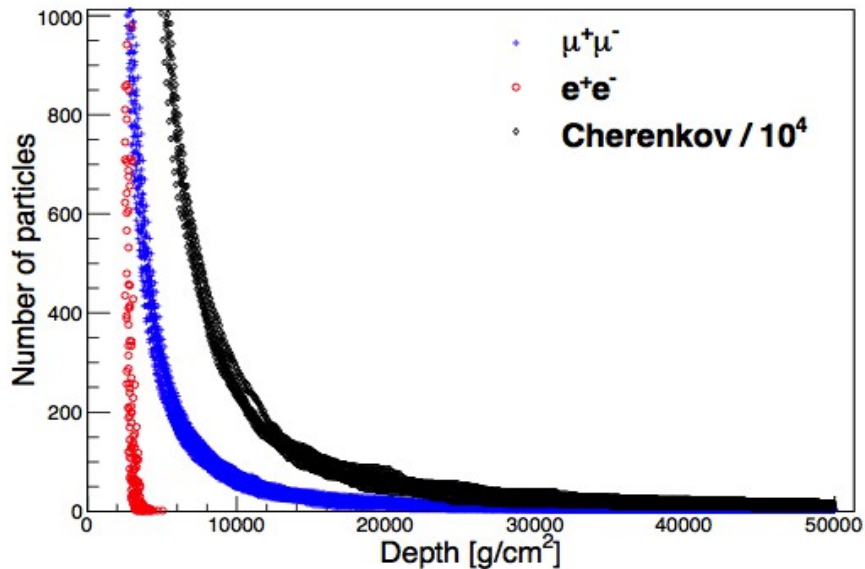
# Back-to-back decay signature @ neutrino observatories



- $QBH \rightarrow \mu^- + p^+$  in water.

$10^{17}$  eV proton +  $3 \times 10^{11}$  eV muon

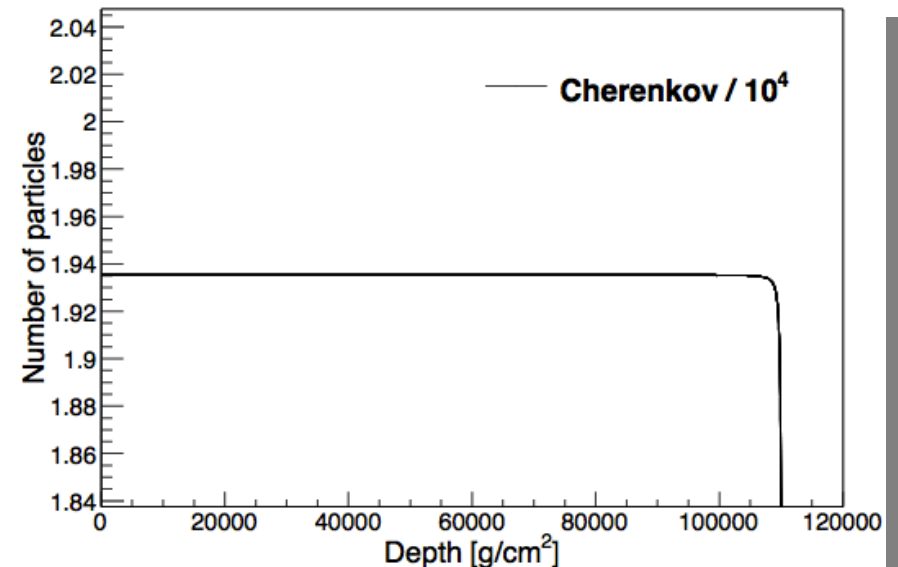
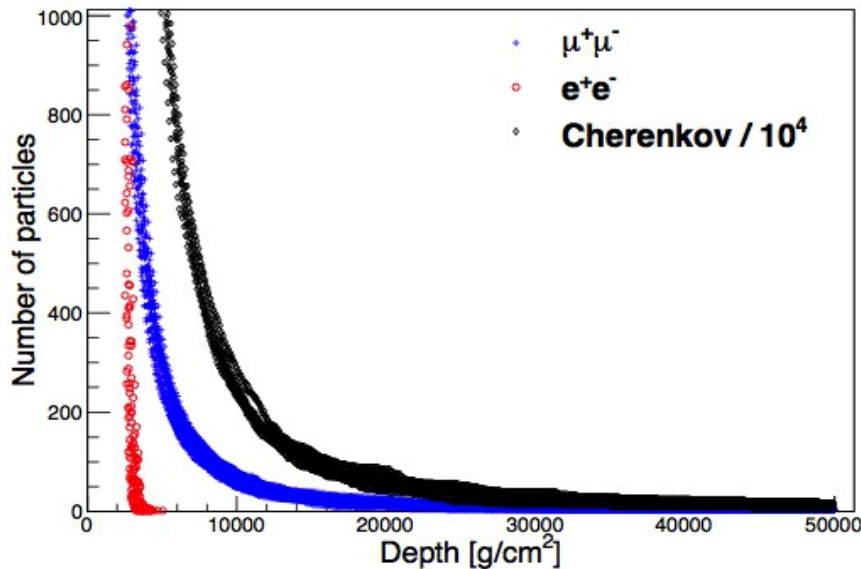
- Angle in the Earth Frame:  $5^\circ$ .
- The EM component of the  $p^+$  shower dies out in about 50 meters!
- Muons propagate for distances of 1- 2 km emitting Cherenkov light!



# Back-to-back decay signature @ neutrino observatories

No. Extra-dimensions	Expected events @ IceCube
1	21
2	81
3	179
4	311
5	440
6	623

- $QBH \rightarrow \mu^- + p^+$  in ICE.
  - Angle in the Earth Frame:  $5^\circ$ .
  - The EM component of the  $p^+$  shower dies out in about 50 meters!
  - Muons propagate for distances of 1- 2 km emitting Cherenkov light!
- The picture looks identical in Ice!



- The back-to-back decay signature of Quantum Black Holes has some unique features:
  - ✓ Two simultaneous particle showers oriented at an angle and pointing to a common origin;
  - ✓ The reconstructed energies for the showers have specific values!
- With this there are two more kinds of experiments which can look for TeV scale black holes and in turn for the Planck scale:
  - @ cosmic rays experiments can look for higher mass values
    - Earth based experiments (Pierre Auger Observatory)
    - Future Space observatories (JEM-EUSO will have an acceptance 30 times larger than the Pierre Auger Observatory)
  - @ neutrino observatories
    - IceCube experiment (running)
    - KM3NeT future neutrino telescope (when running will have a volume several times higher than IceCube).